

Understanding the Variable Sky with Near-simultaneous, Multi-filter Photometry: Stellar Activity and Small Bodies

SEO-WON CHANG (장서원)

Seoul National University



SCIENCE
자연과학대학 · 기초과학연구원



Australian
National
University

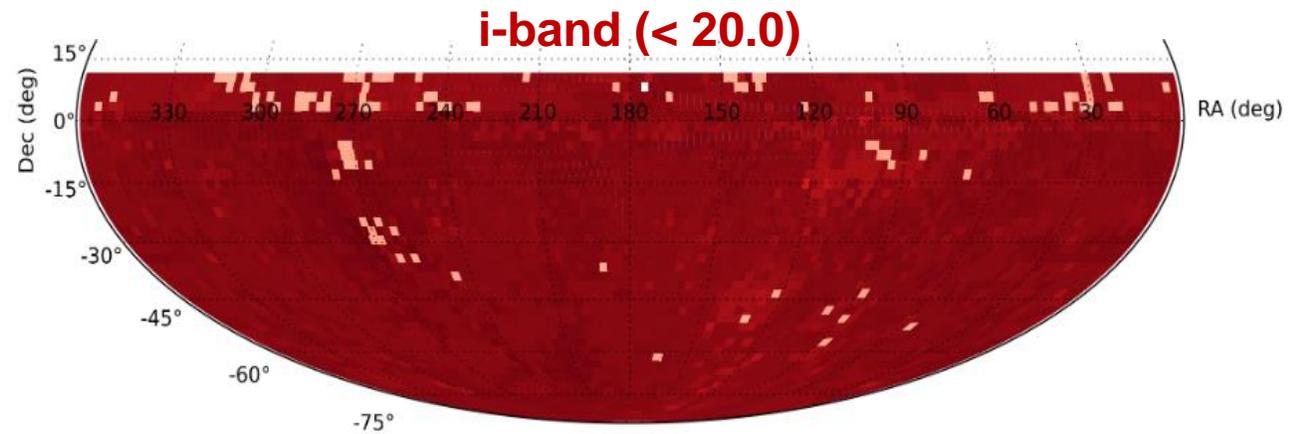
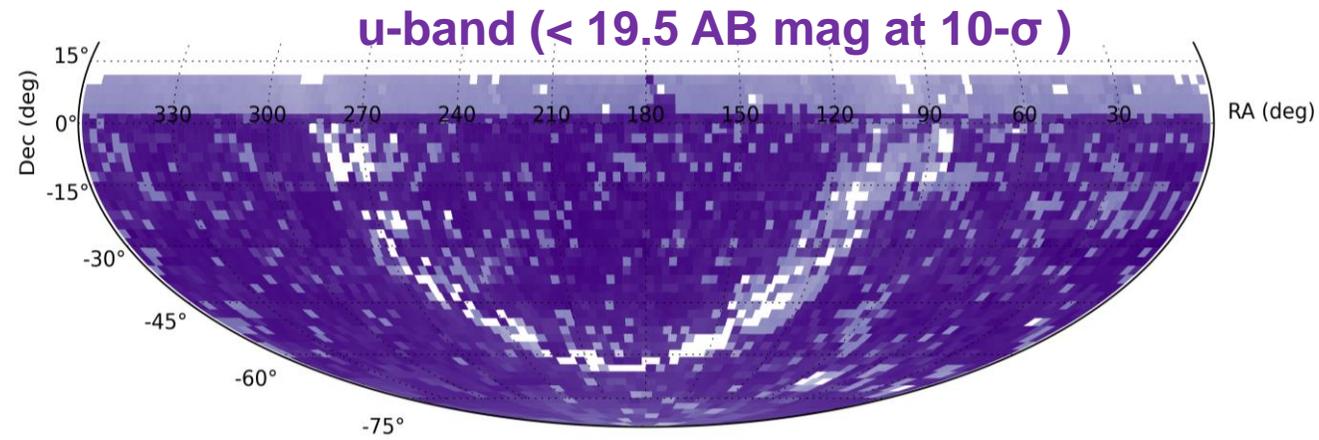
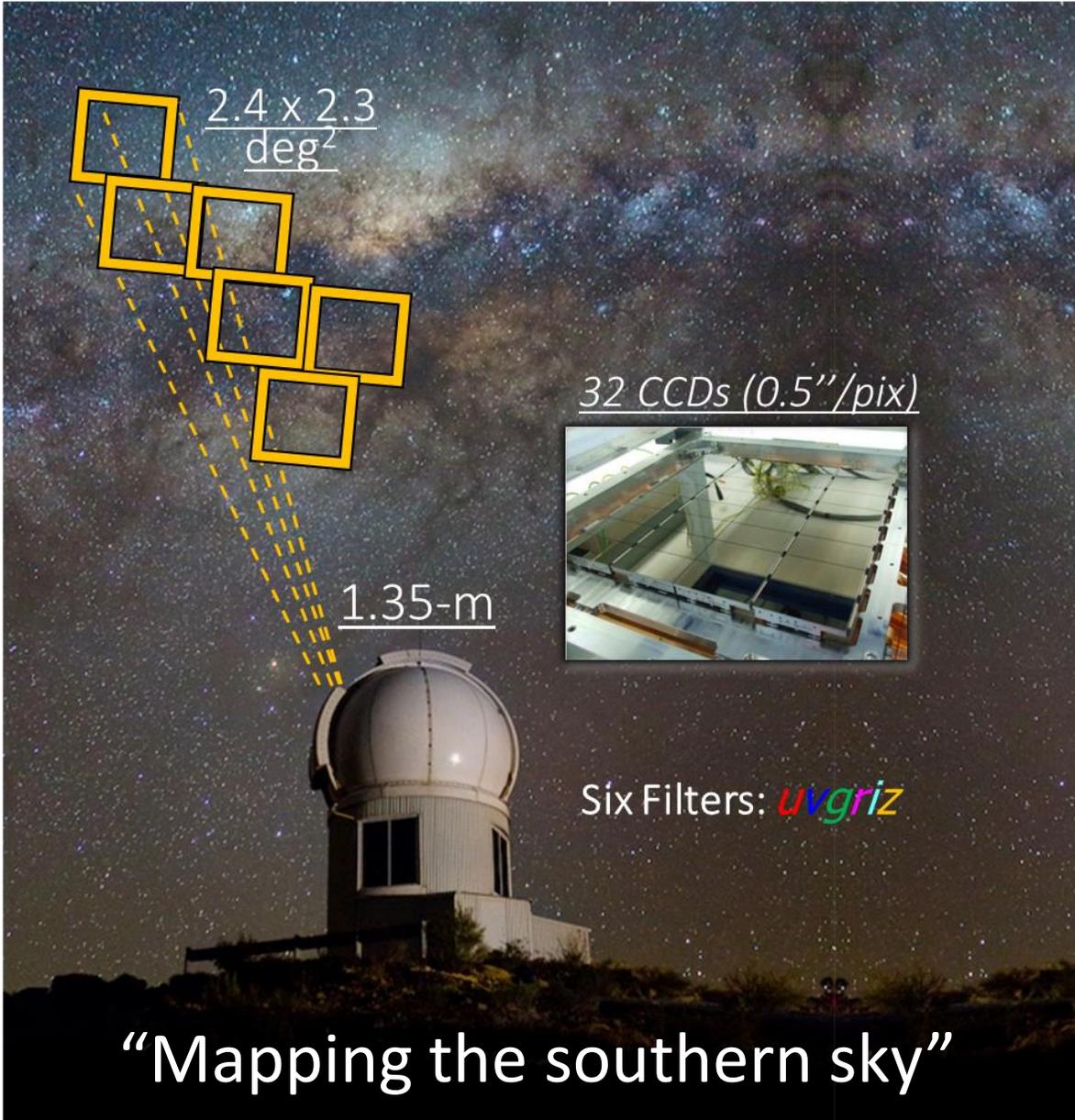


What I'm going to tell you

- SkyMapper Southern Sky Survey: **Near-simultaneous, Multi-filter Photometry**
- Time-Domain Science: **Stellar Activity** and **Small Bodies**
Part of this talk represents results from Chang+ 2020 ([arXiv: 1910.06478](#)) and Sergeyev+ 2021 ([arXiv: 2110.11656](#))
- Recommendation for 7DT Galactic Time-Domain Science

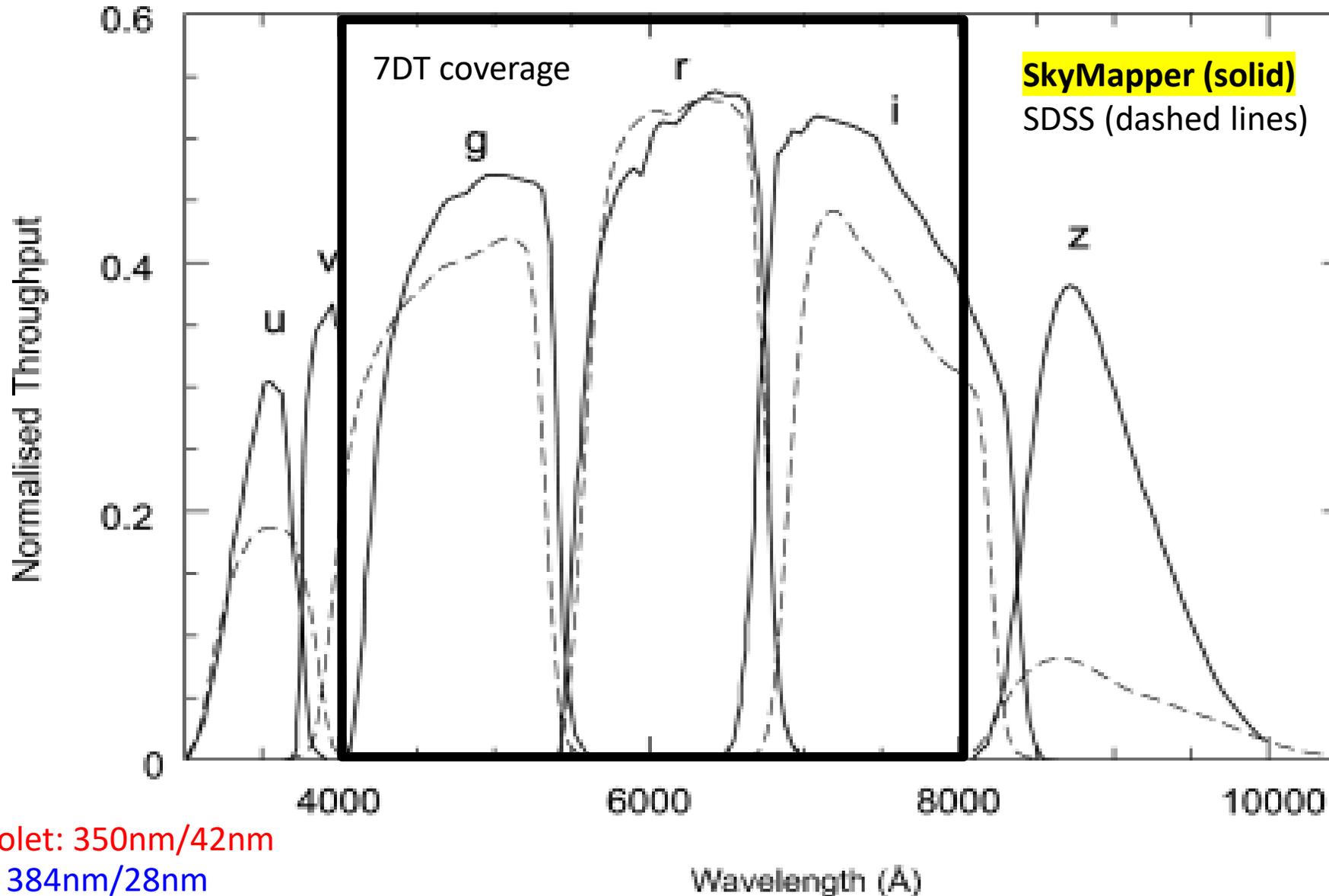
DISCUSS MORE POSSIBLE SCIENCE CASES AND ENGAGE DOMESTIC SCIENCE TEAMS FOR FEEDBACK, NOT JUST INTERNAL CORE COLLABORATION

Intro: SkyMapper Southern Sky Survey



DR3 Historic data
(March 2014 to October 2019)

Broadband Filters: *uvgriz*

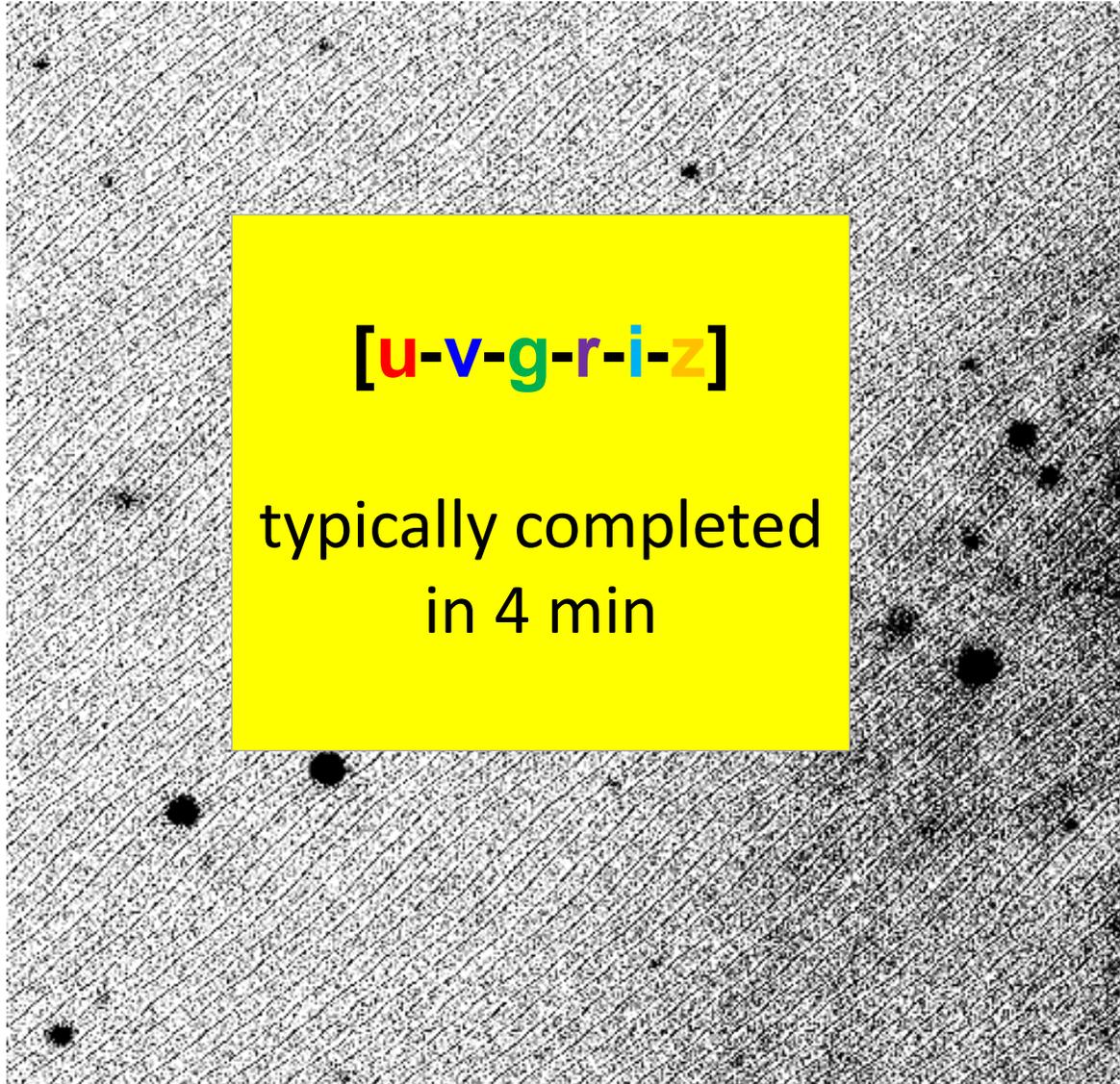


Ultraviolet: 350nm/42nm

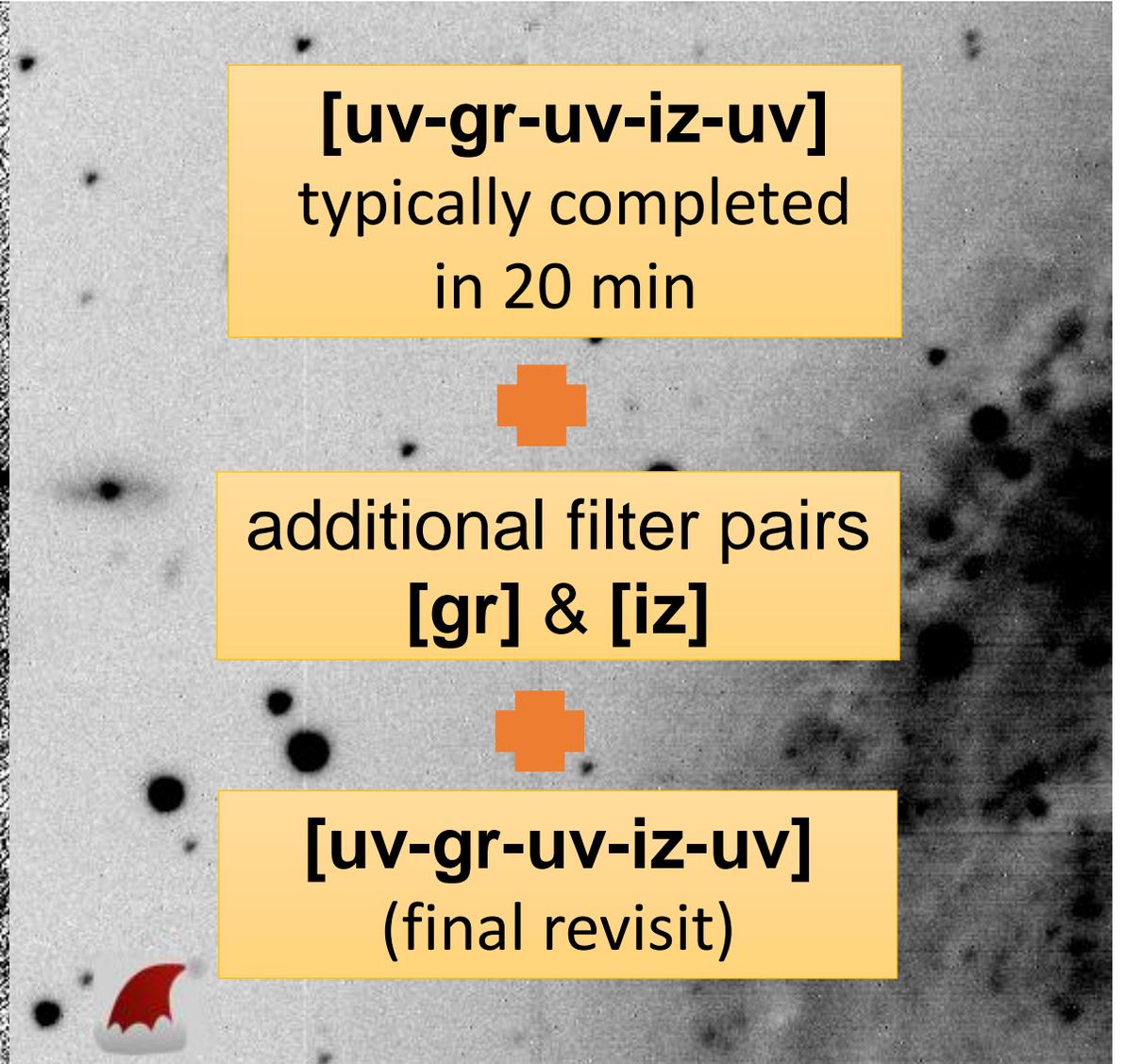
Violet: 384nm/28nm

Sensitive to
stellar
atmospheric
parameters ~
(T_{eff} , $\log g$ and
[Fe/H])

Near-simultaneous, Multi-filter Photometry Data



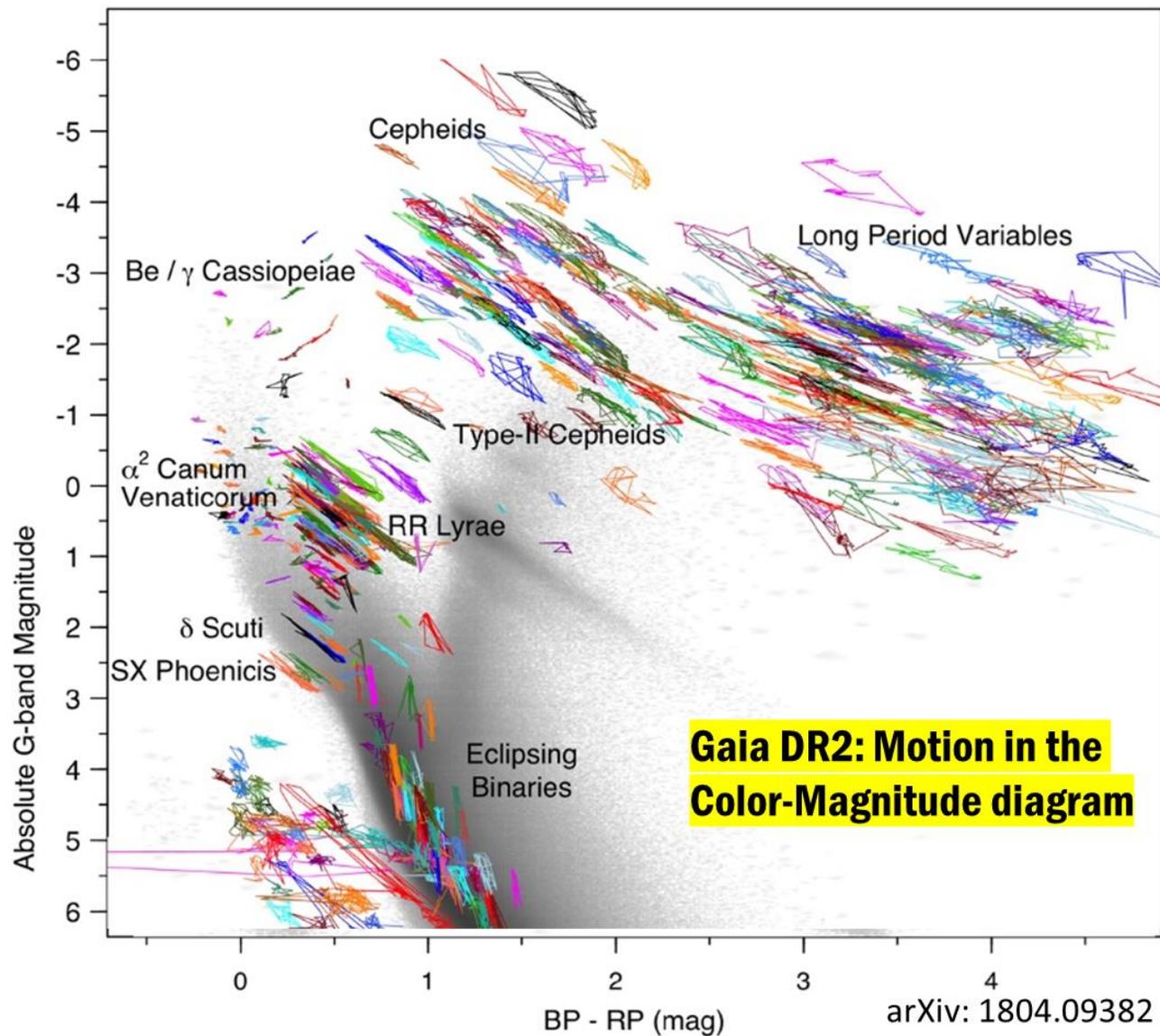
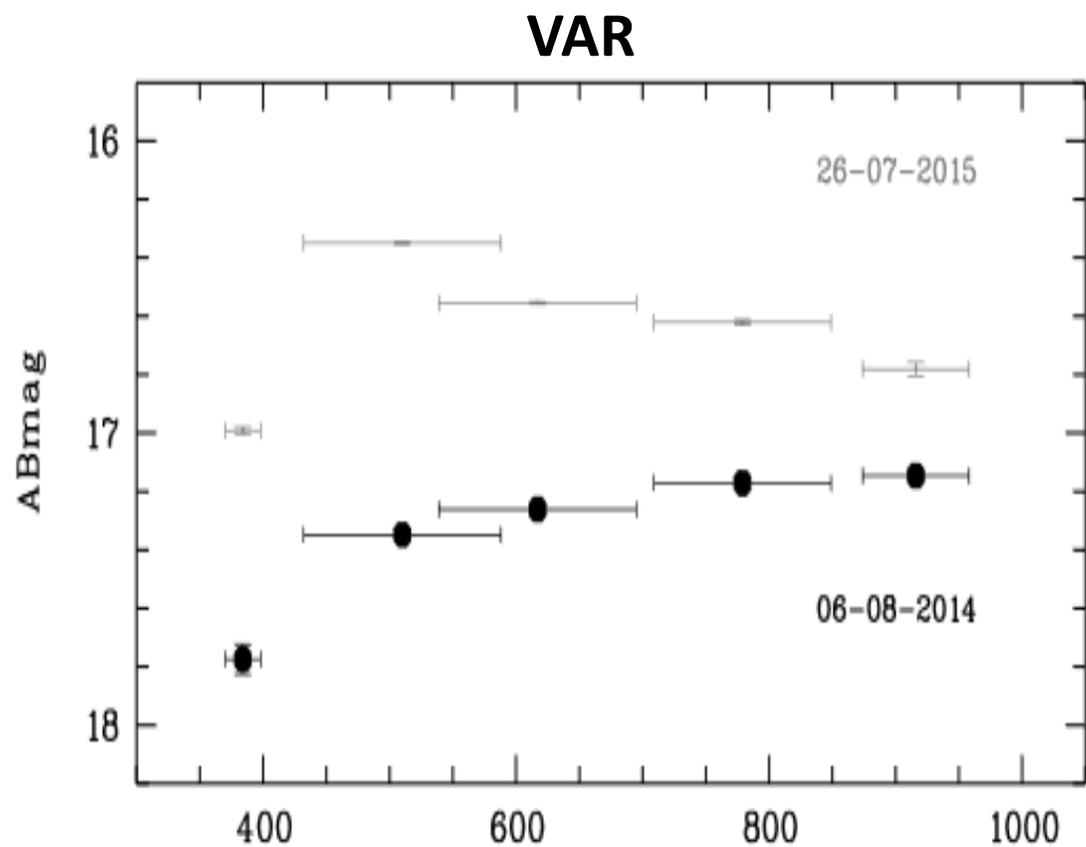
Shallow Survey (9 to 18 mag in all six bands)



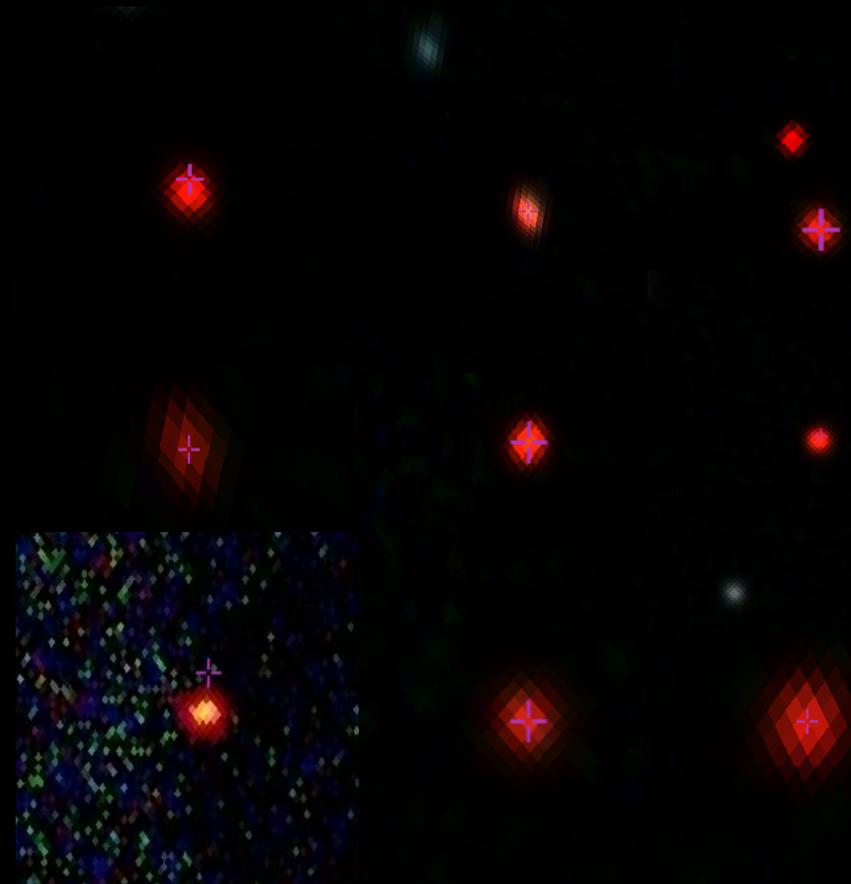
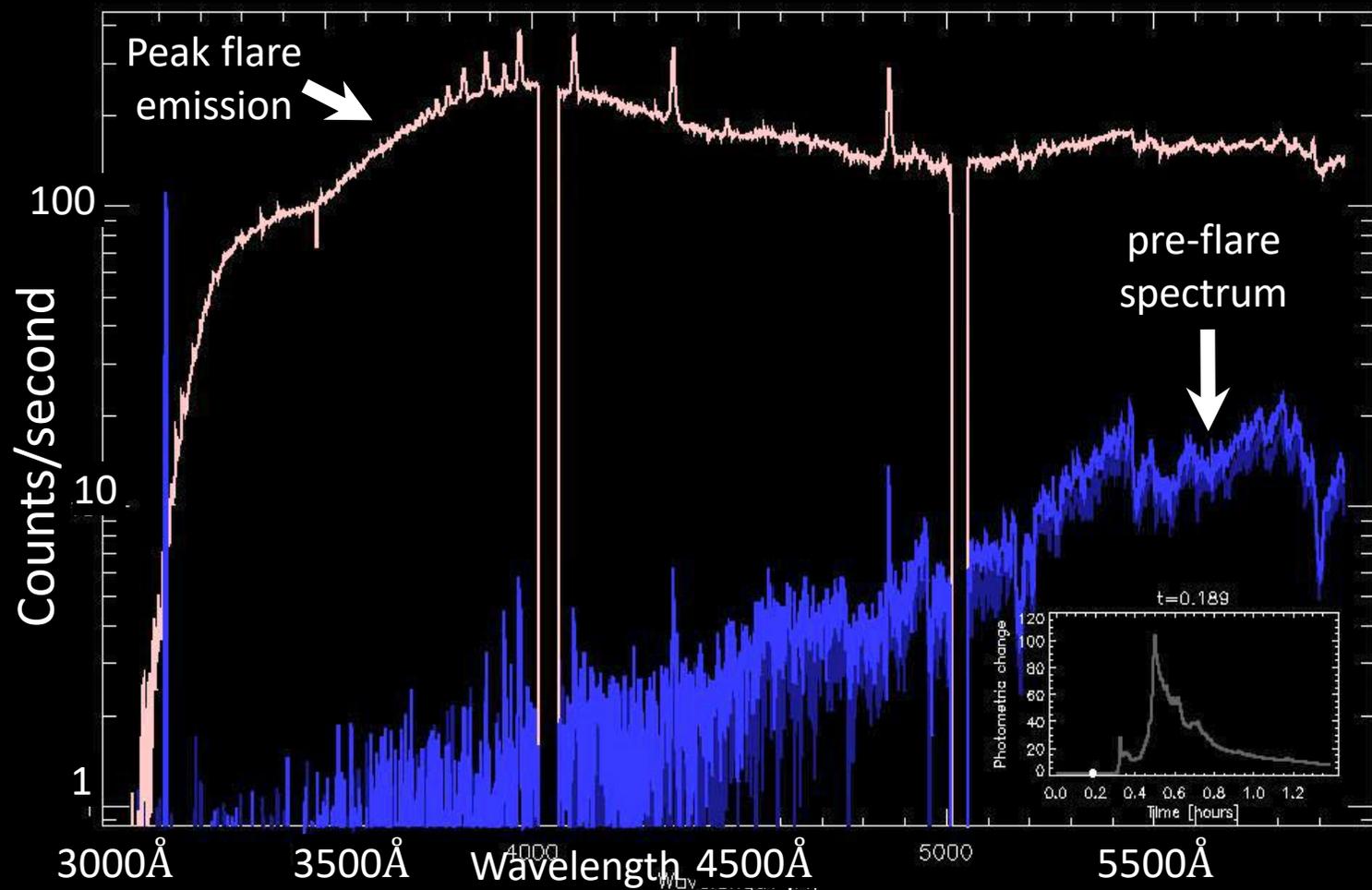
Main Survey (Down to 21 mag)

Time-dependent Spectral Energy Distribution (SED)

From SkyMapper's six-filter uvgriz photometry

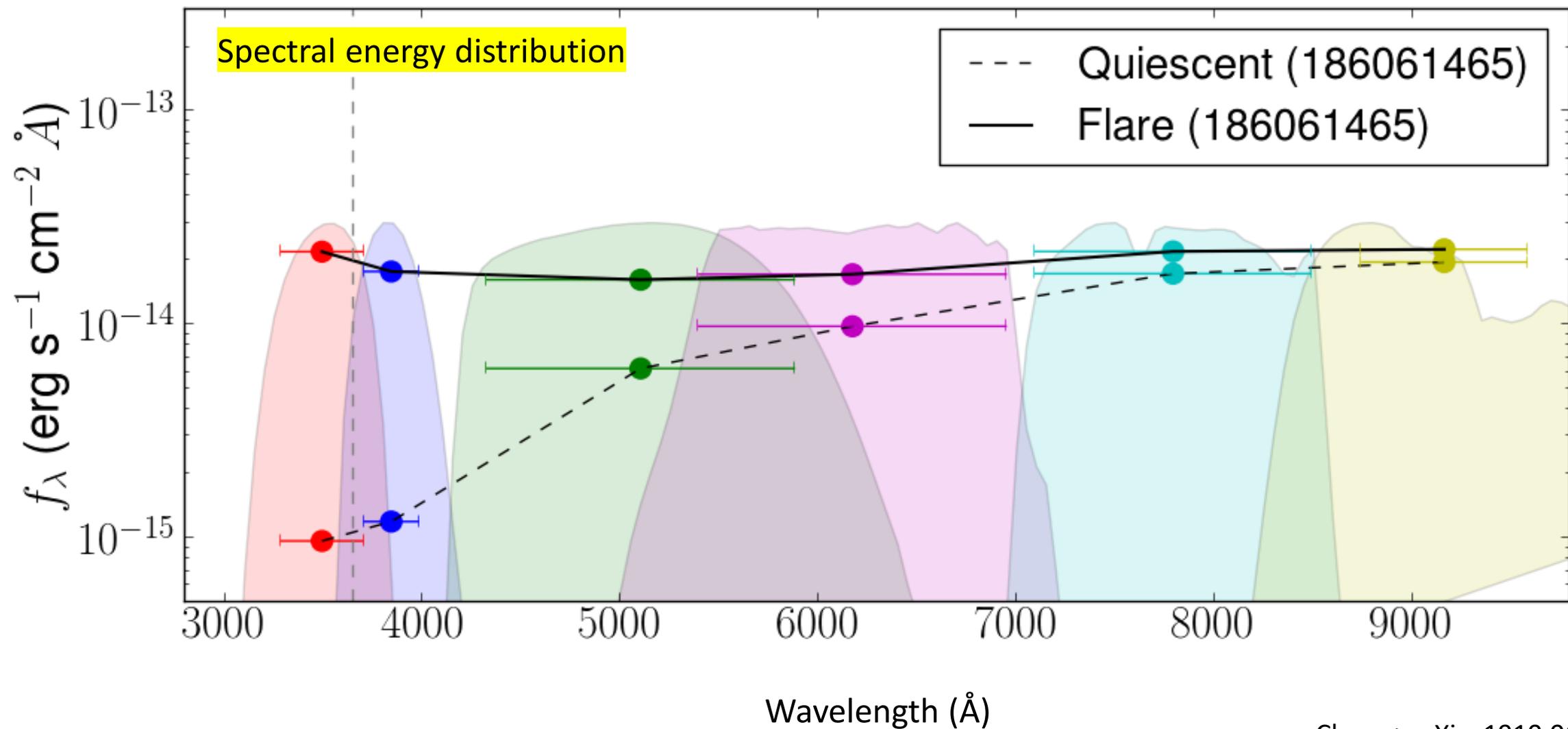


Latest Galactic Sciences: I. Magnetic Activity of **M dwarf**

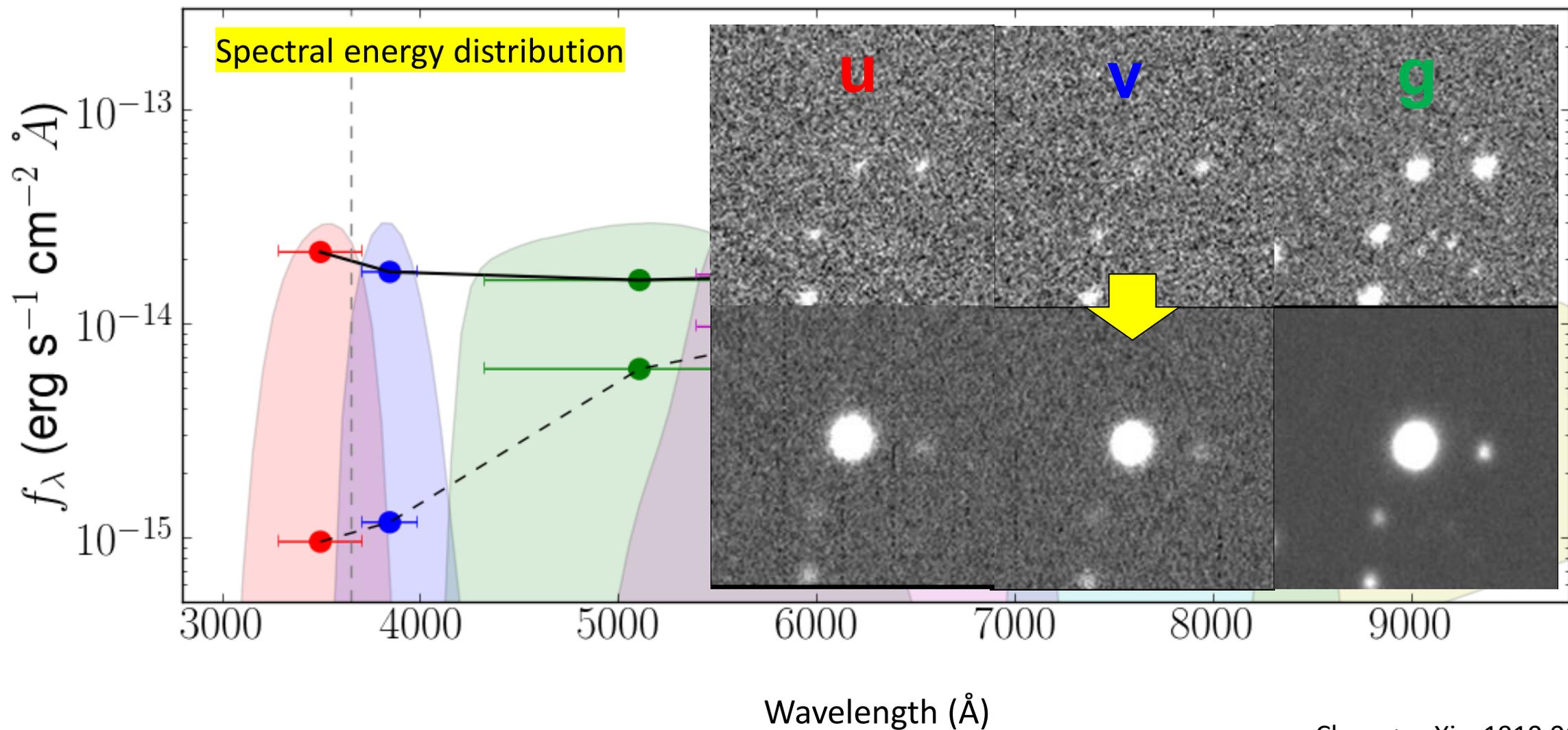


M-dwarf star (YZ CMi), 15 sec cadence, $R \sim 1000$, 70x faster than VLT! (obs Jan 2012, Spectra: Brown et al. SALT/RSS; photometry: Kowalski et al. WHT/La Palma)

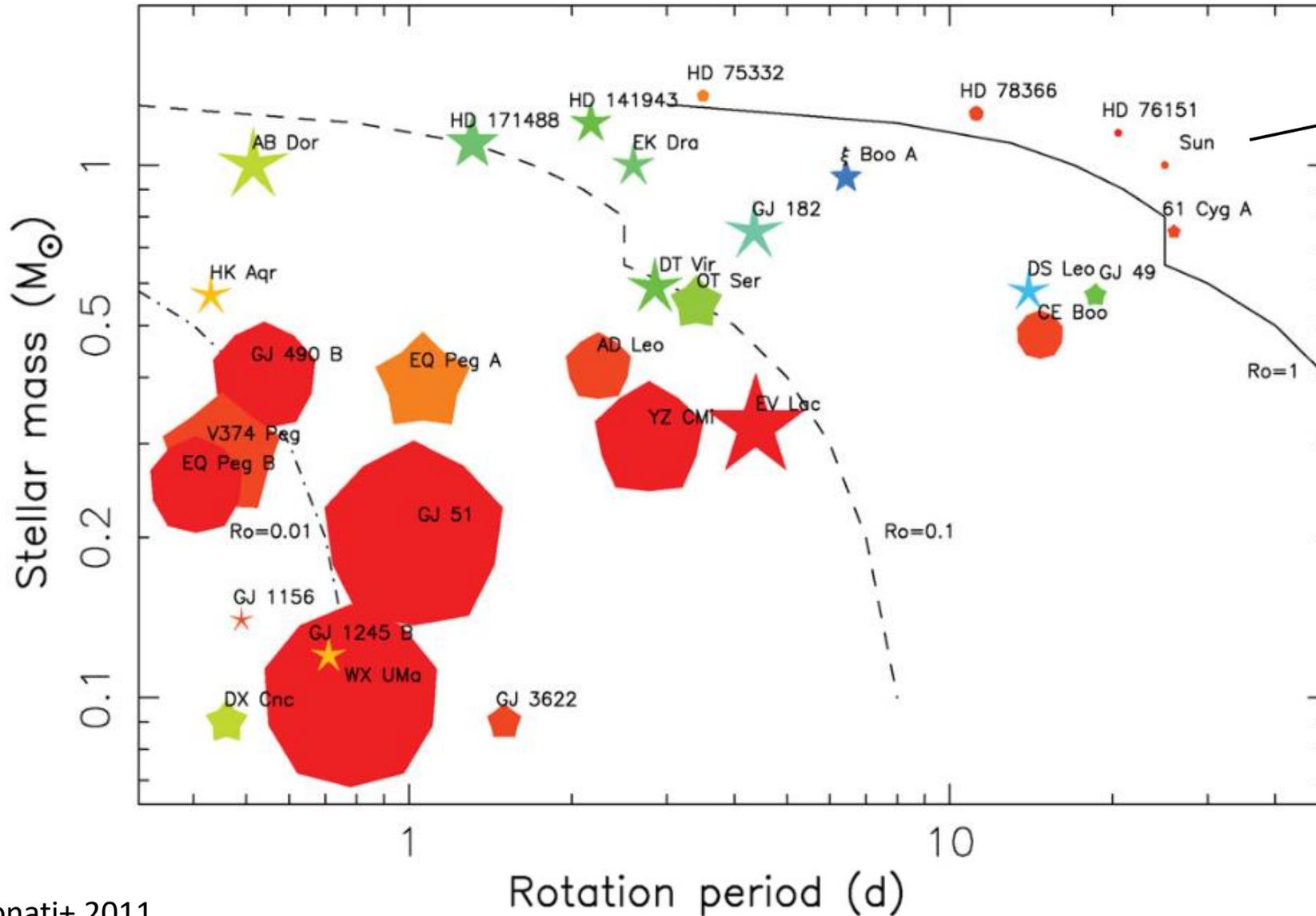
Latest Galactic Sciences: I. Magnetic Activity of **M dwarf**



Latest Galactic Sciences: I. Magnetic Activity of **M dwarf**



Magnetic Fields of M Dwarfs



* Point size
= strength of
magnetic fields

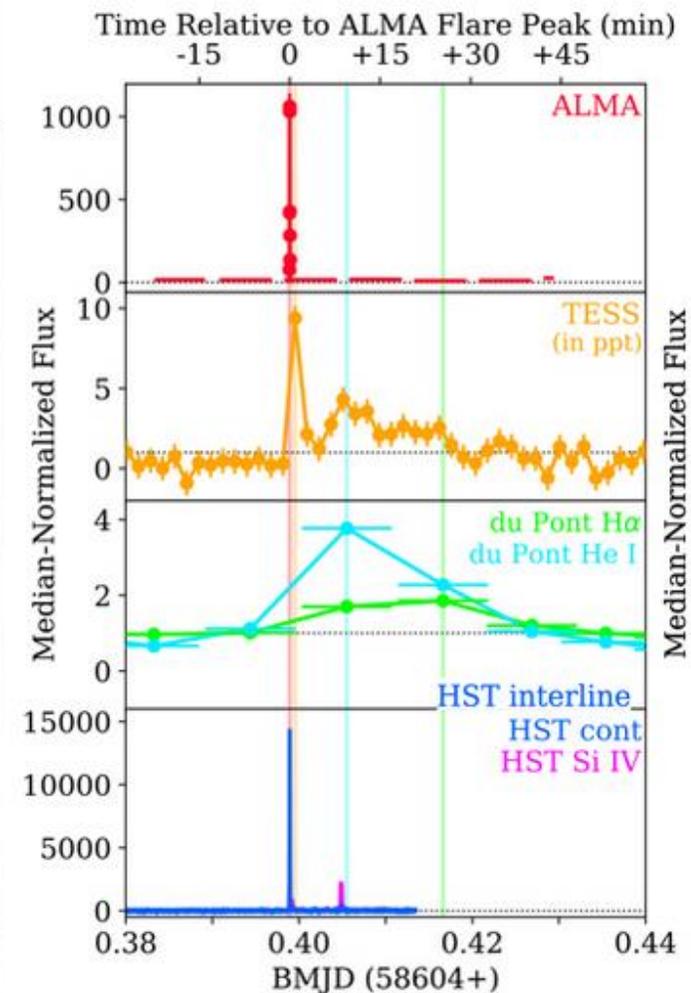
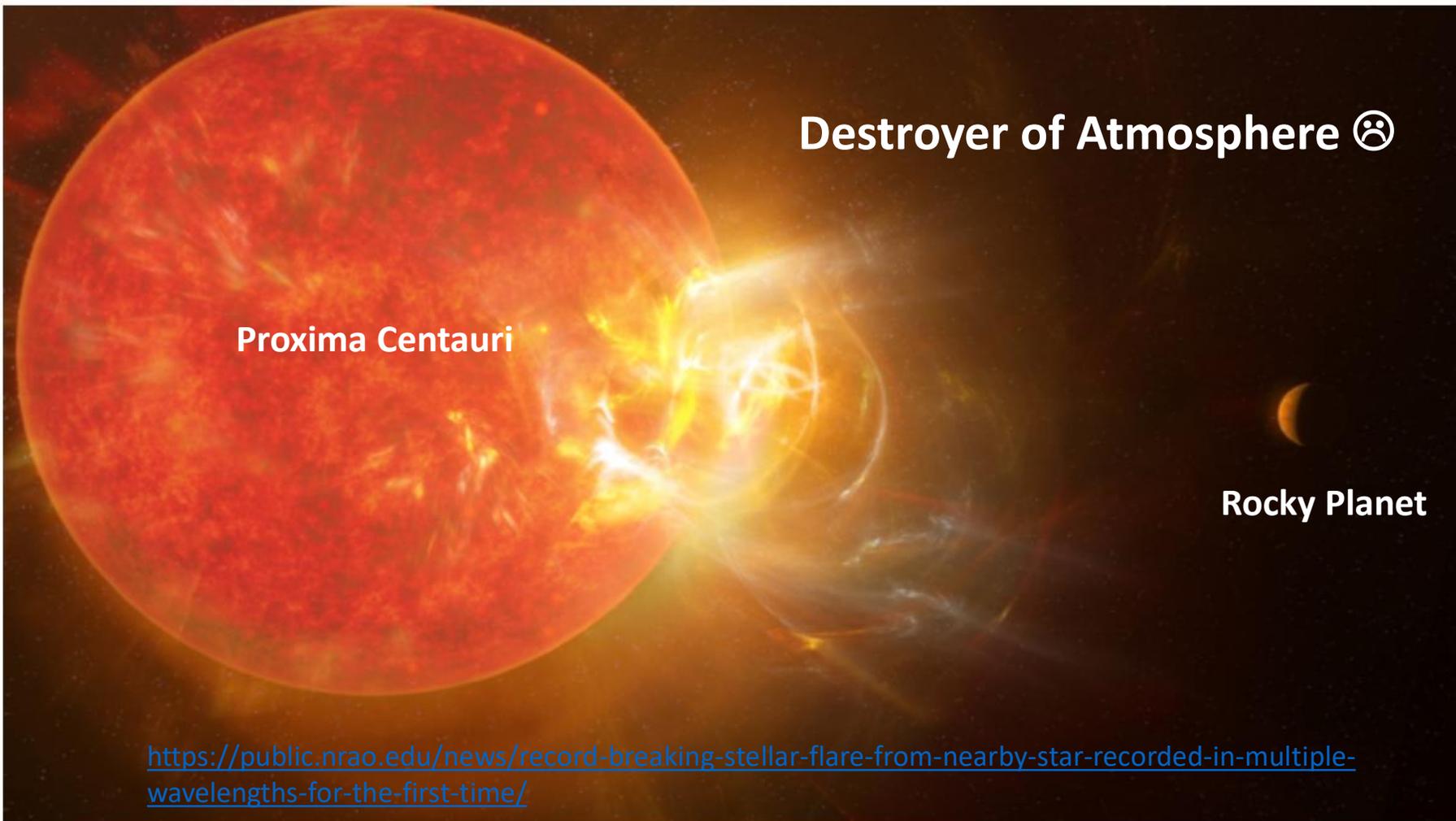
**Strongly
magnetic
objects
(frequently
exceed 1 kG)**

Vs.

Sun (10-50 G)

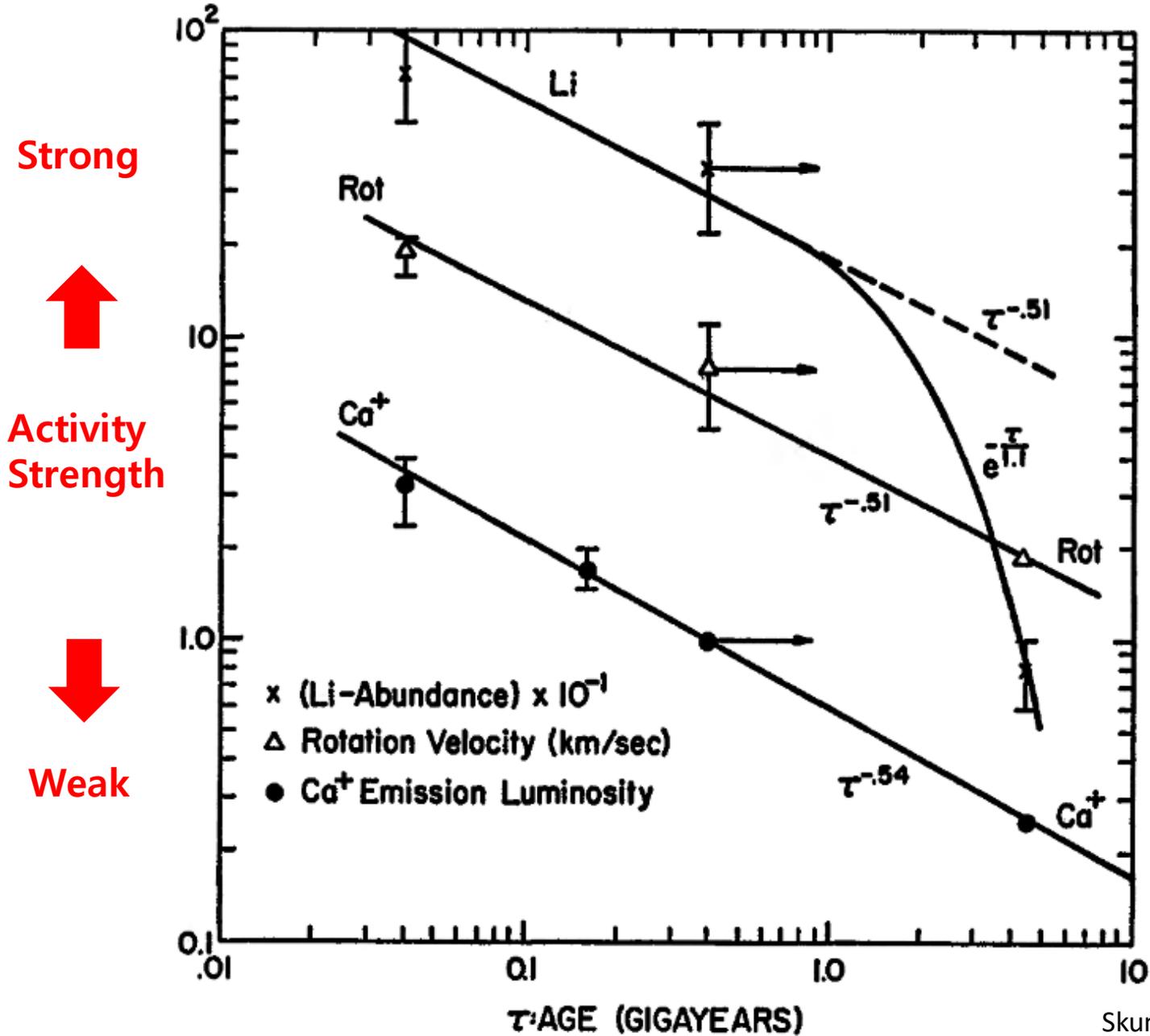
Home > News > News Release: April 21, 2021 at 10:00 am EDT

Record-breaking stellar flare from nearby star recorded in multiple wavelengths for the first time



2019 May 1 event at all wavelengths (MacGregor+ arXiv: 2104.09519)

Skumanich Law: Age-Rotation-Activity Relation



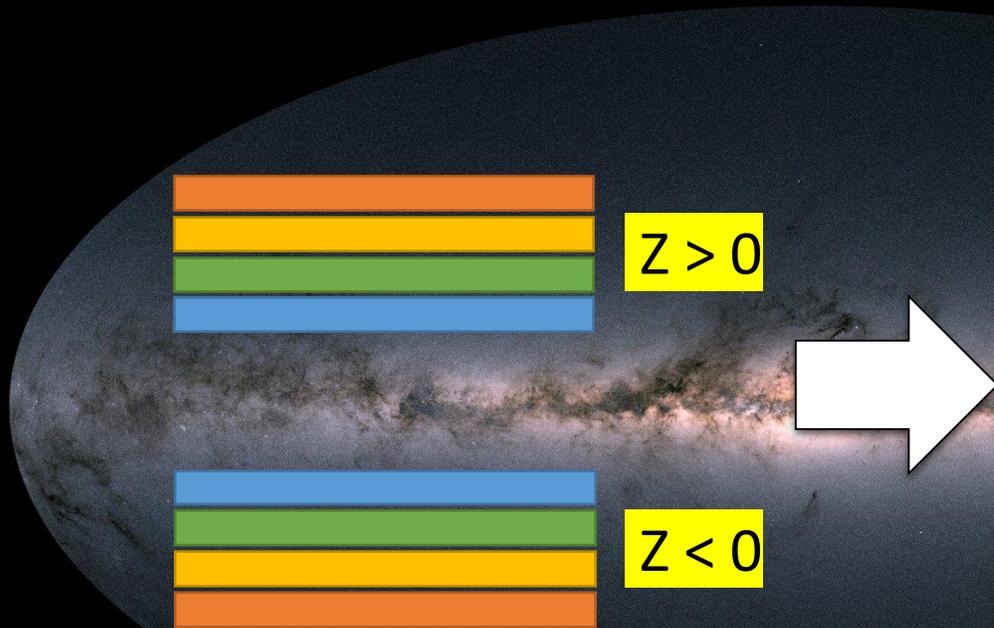
Strong
 ↑
 Activity Strength
 ↓
 Weak



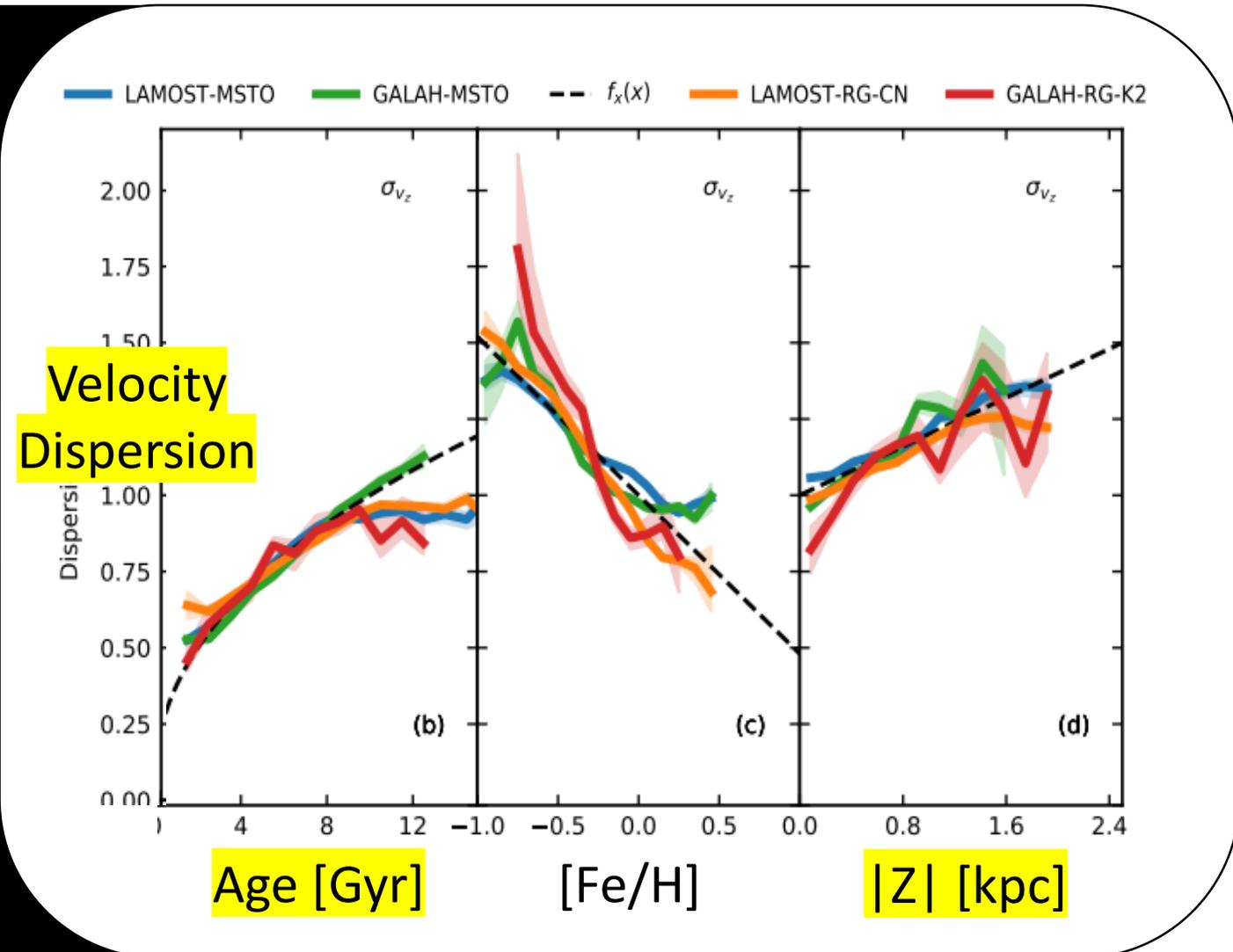
← Stellar rotation

Young ← Age → Old

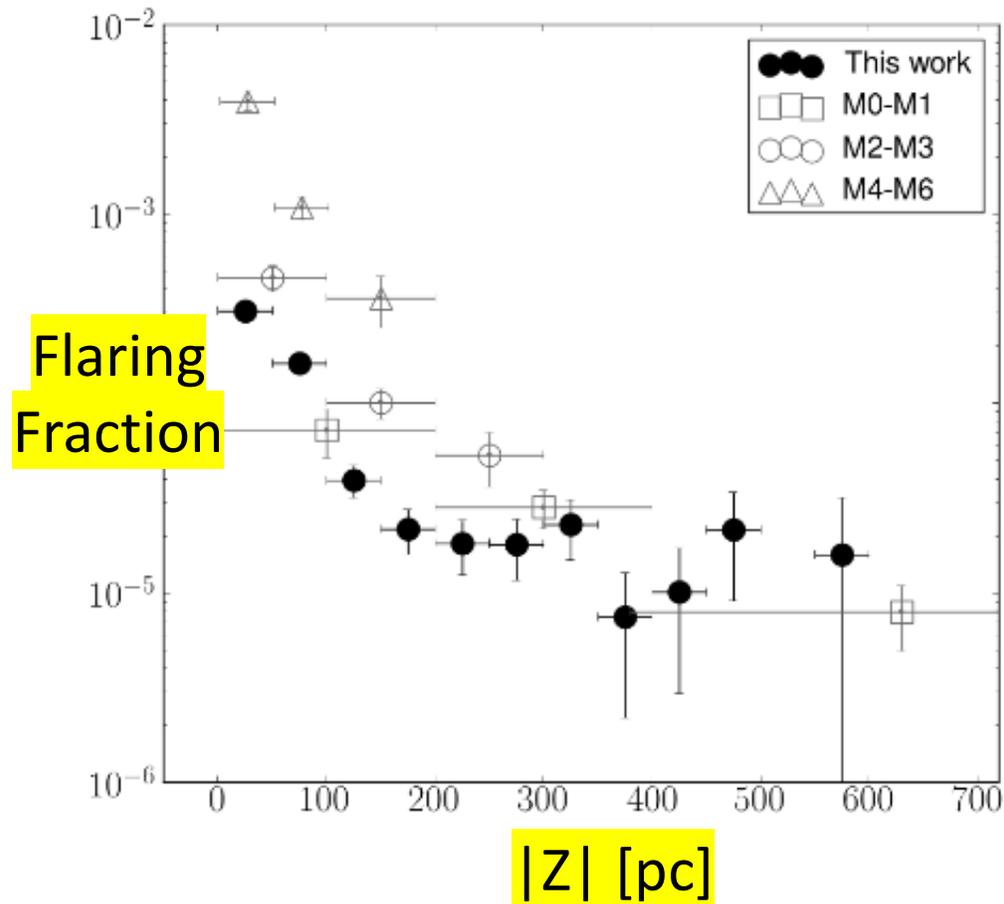
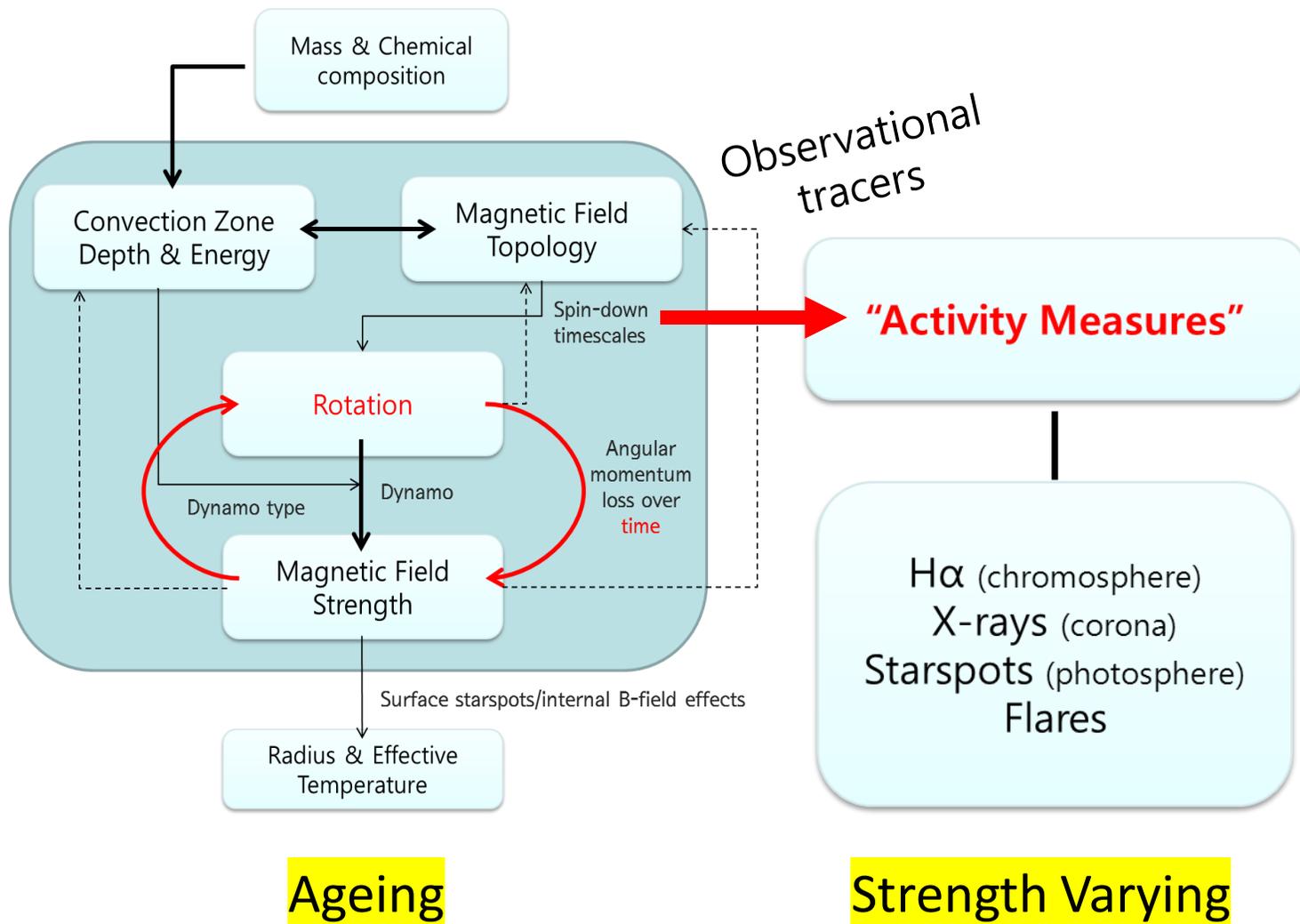
Vertical AGE gradient in the Milky Way disc: Dynamical Evolution



Credit: ESA/Gaia/DPAC

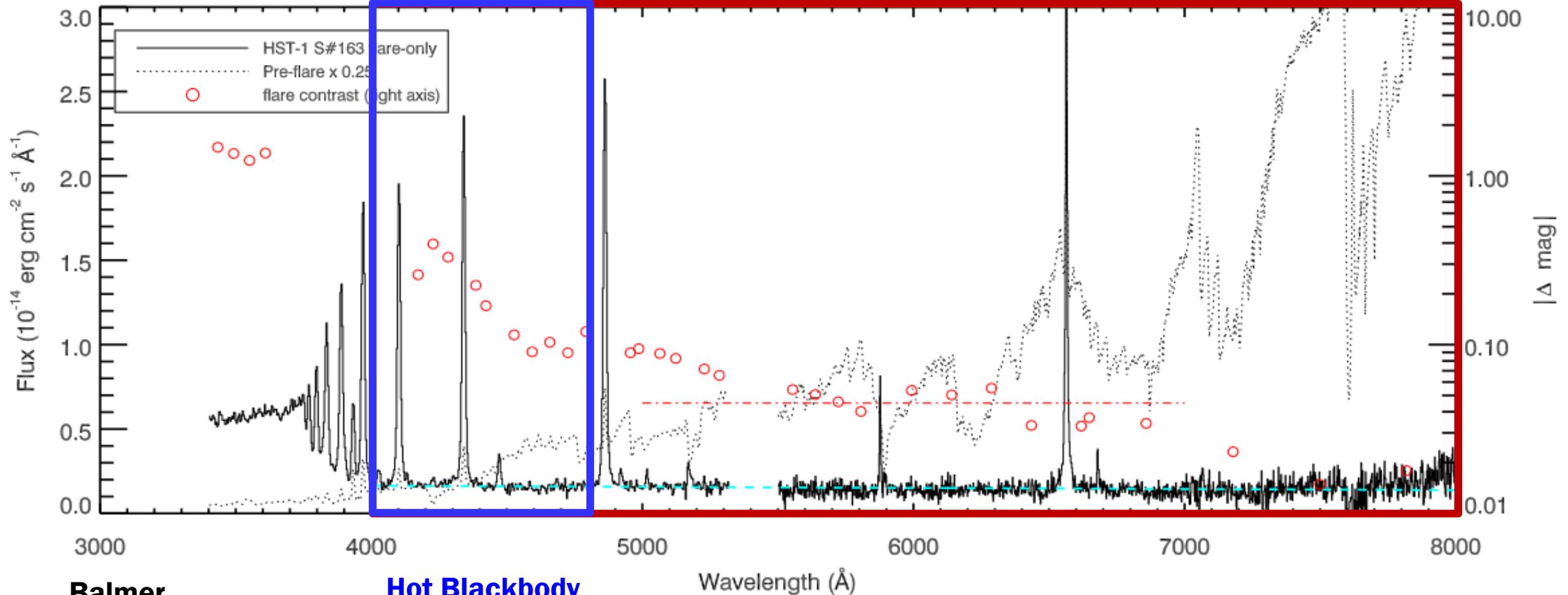


Another Proxy for AGE



Two- or Three-component flare model

7DT coverage



**Balmer
Continuum
Component:
< 3,647Å**

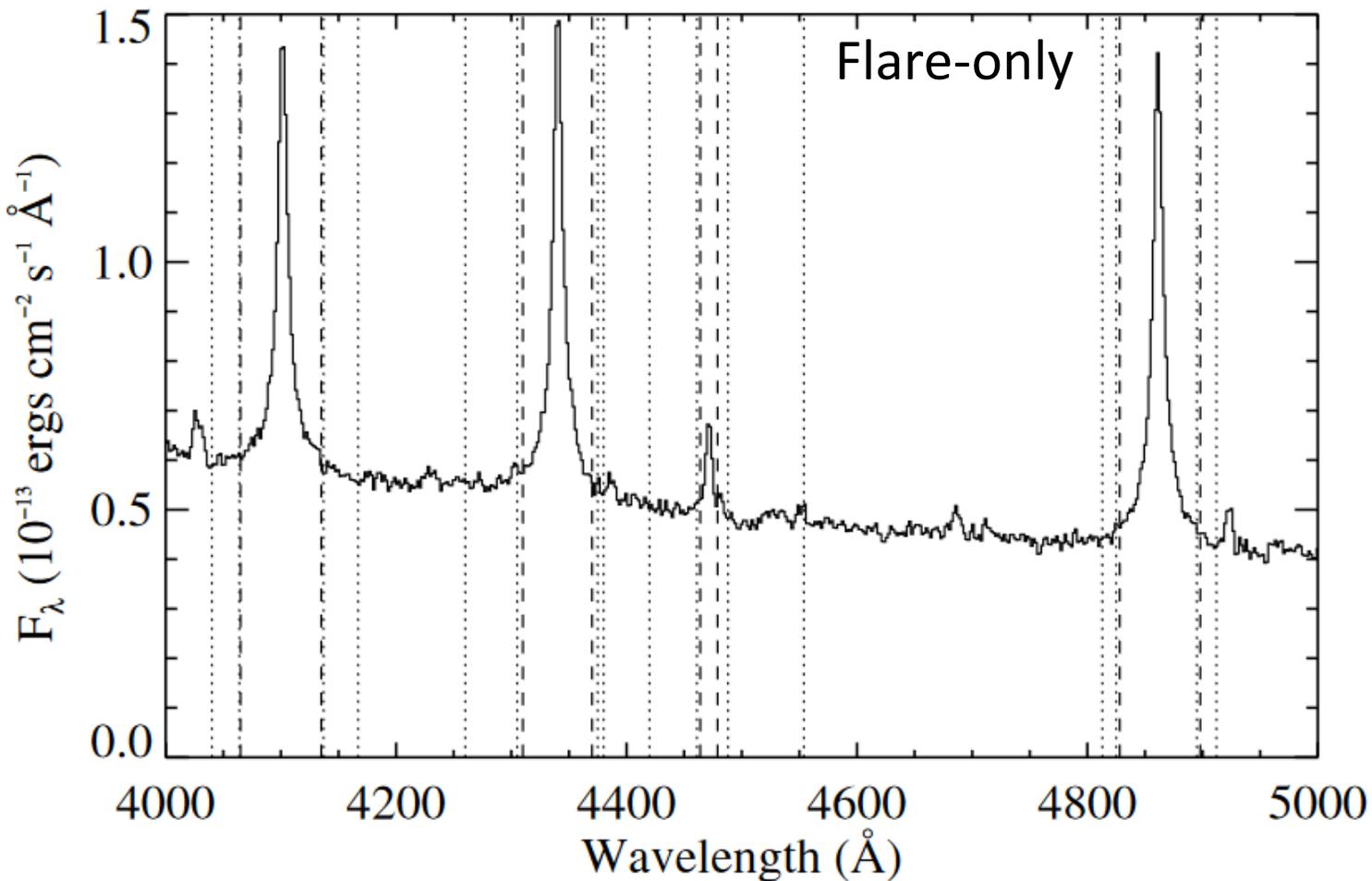


**Hot Blackbody
Component
(~10,000K;
4,000 ~ 4,800Å)**



**Redder continuum:
> 4,900Å**

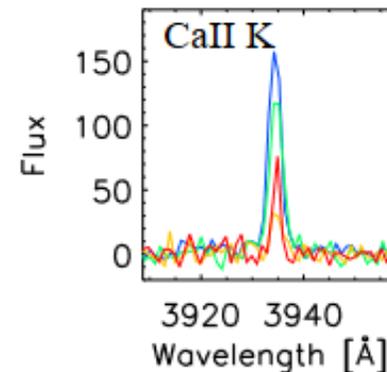
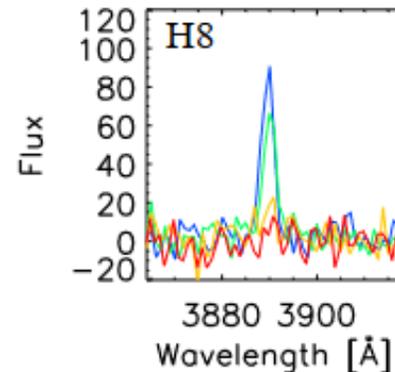
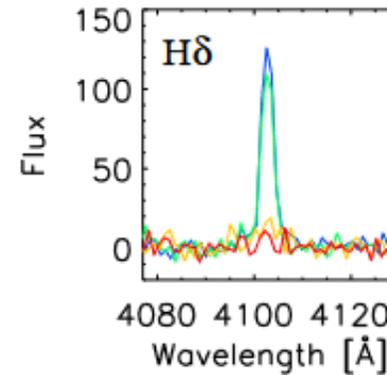
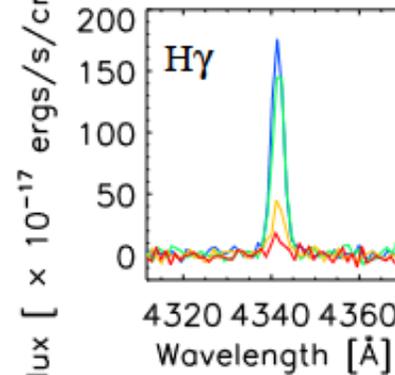
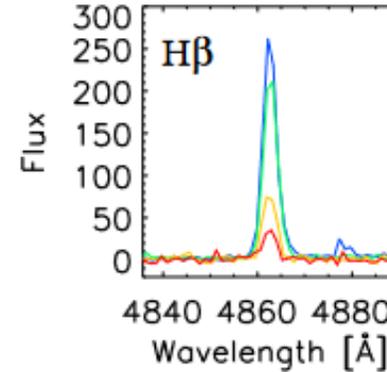
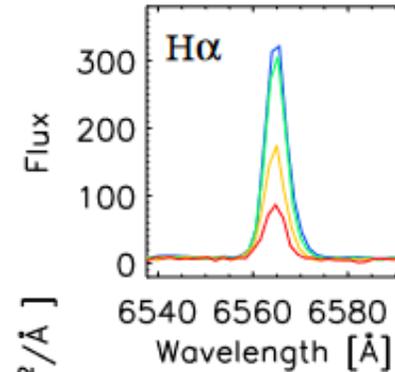
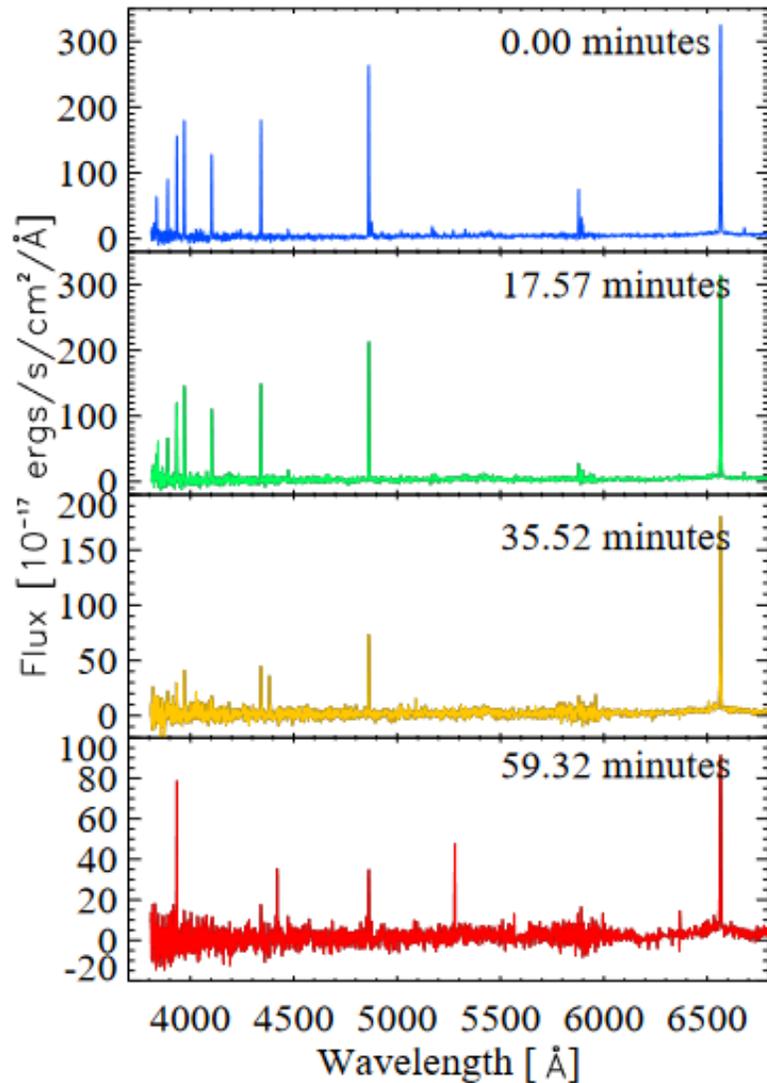
Key Emission Lines: hydrogen Balmer, Helium I, ...



Line	Line Integration Window (Å)
H α	6520–6610
H β	4828–4898
H γ	4310–4370
H γ	4321–4361
H δ	4065–4135
H δ	4084–4122
He I λ 4471	4464.6–4481
He I λ 4471	4460–4495

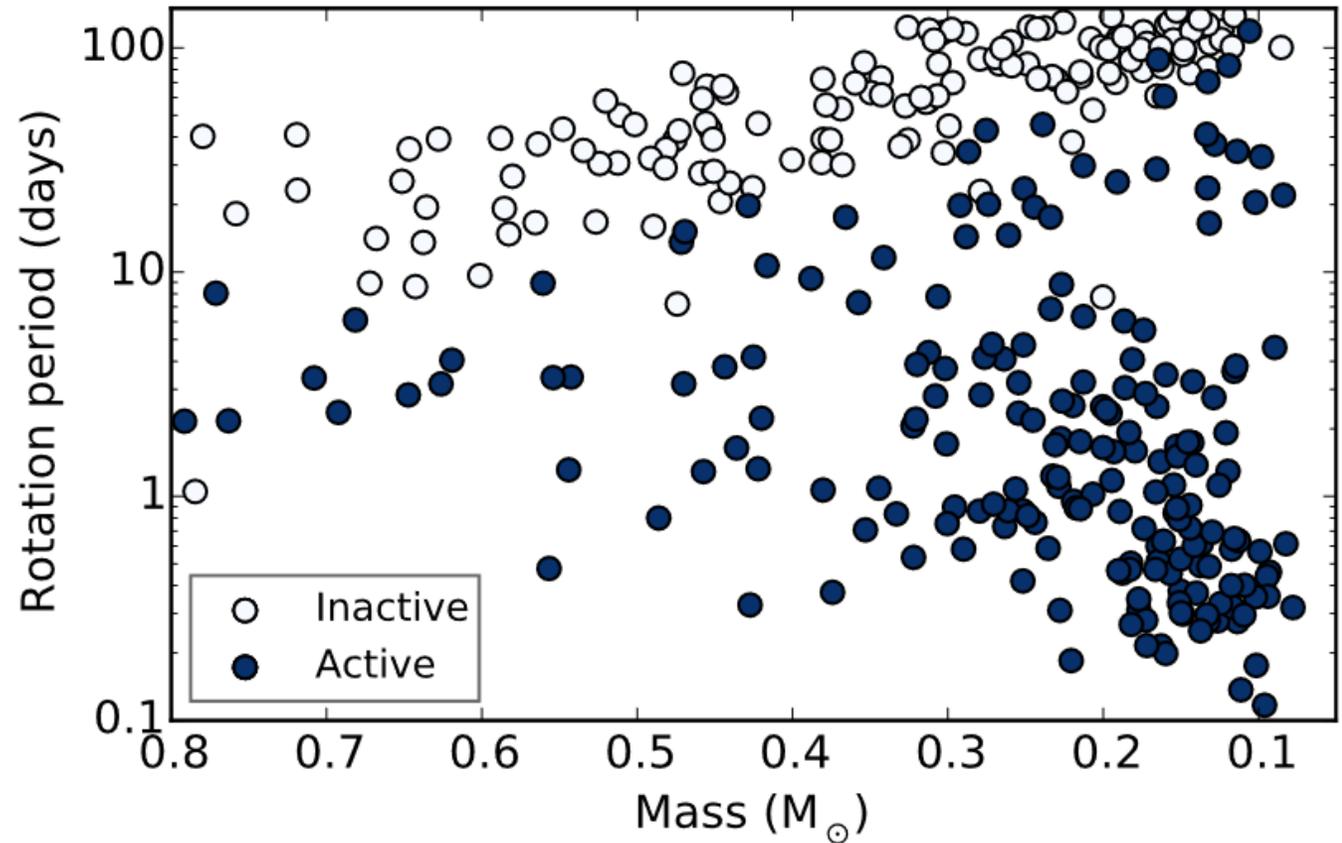
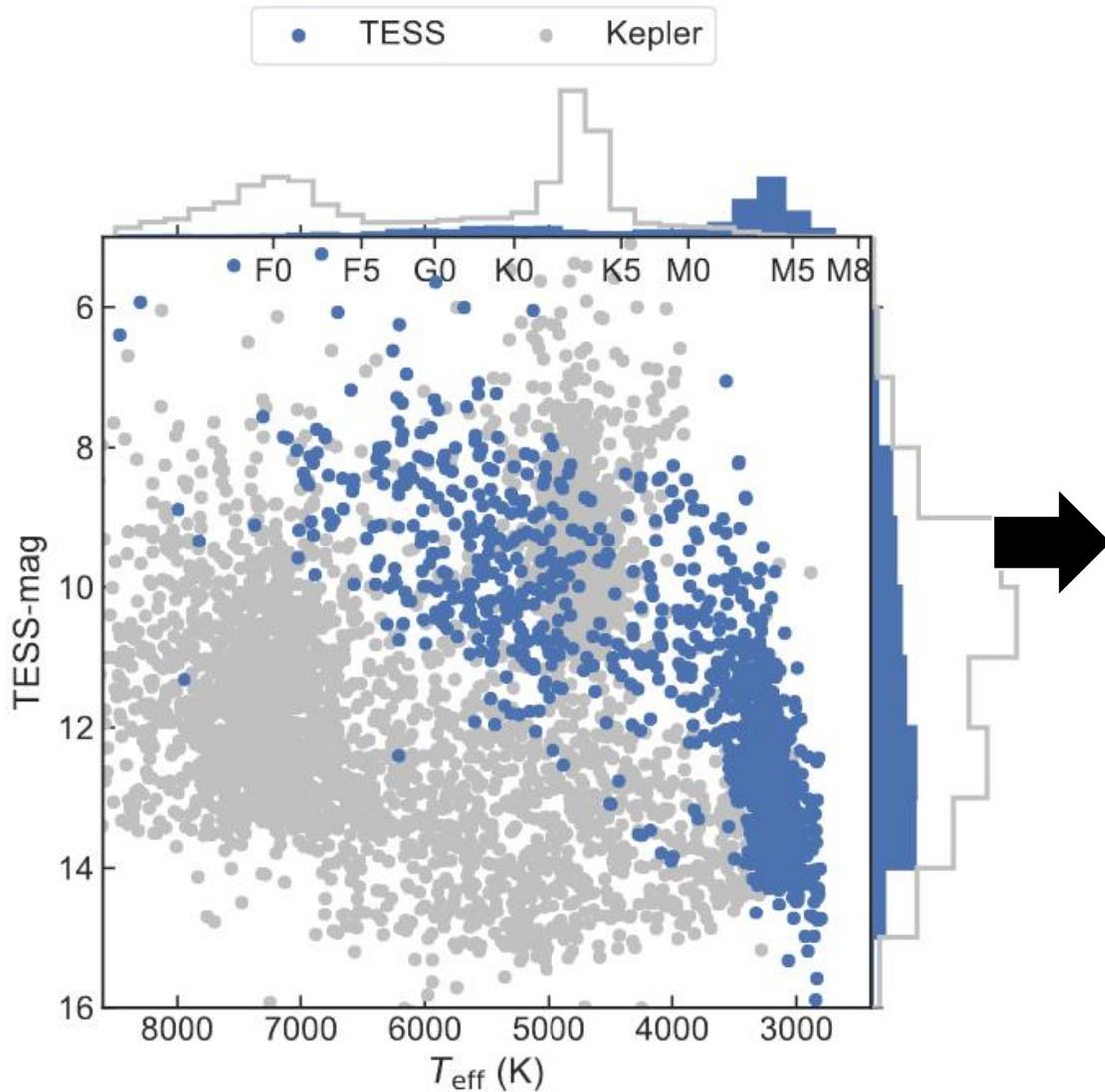
**= 5-12 7DT medium-band filters for
line and continuum measures**

Homogeneous analysis of line and continuum emissions?



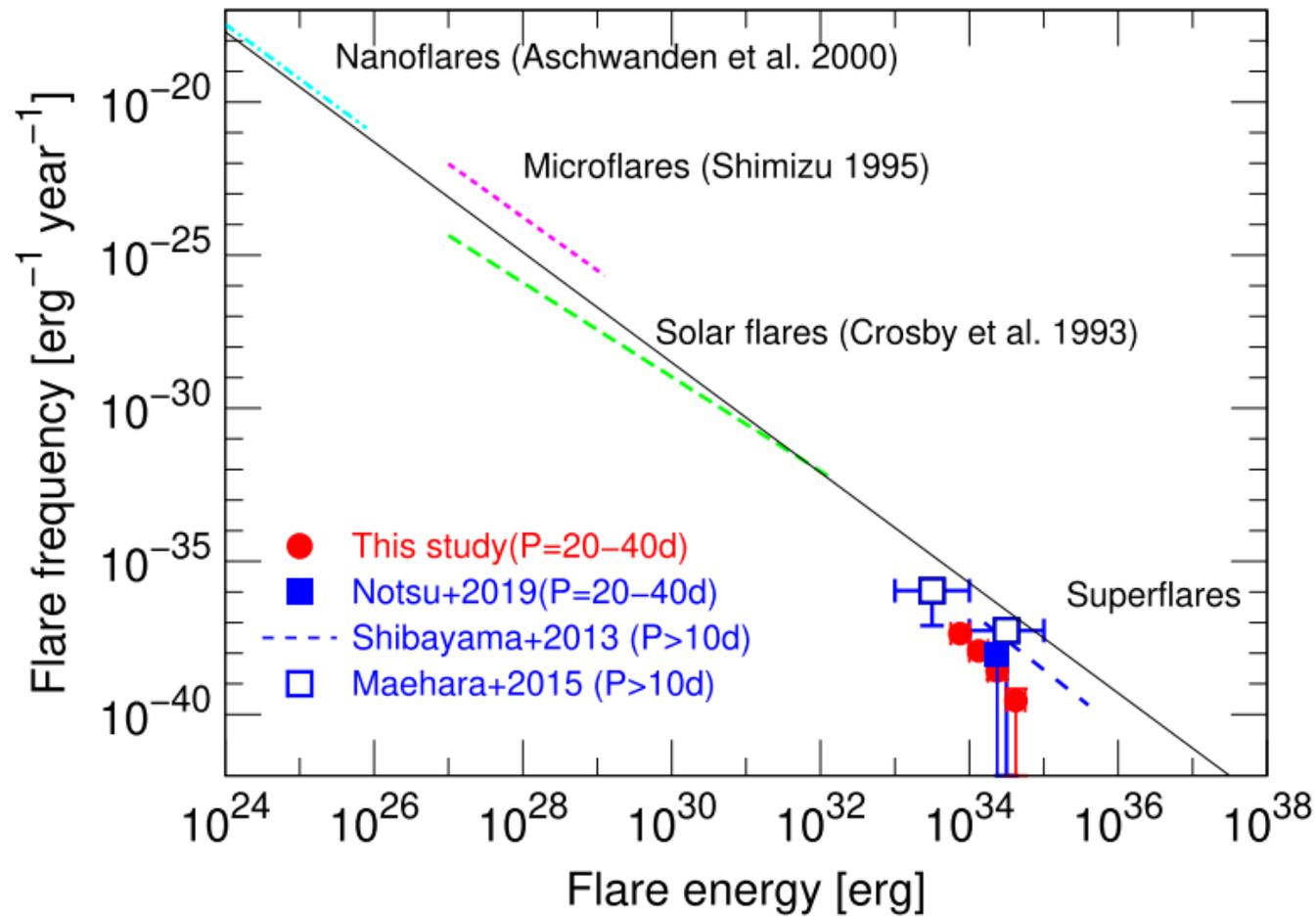
- Previously, the entire Balmer series and several prominent lines have not been captured simultaneously
- Important diagnostic for the evolution of electron densities
- And to break the degeneracy between emission mechanisms

Further constrain the Chromospheric activity of Planet hosts?

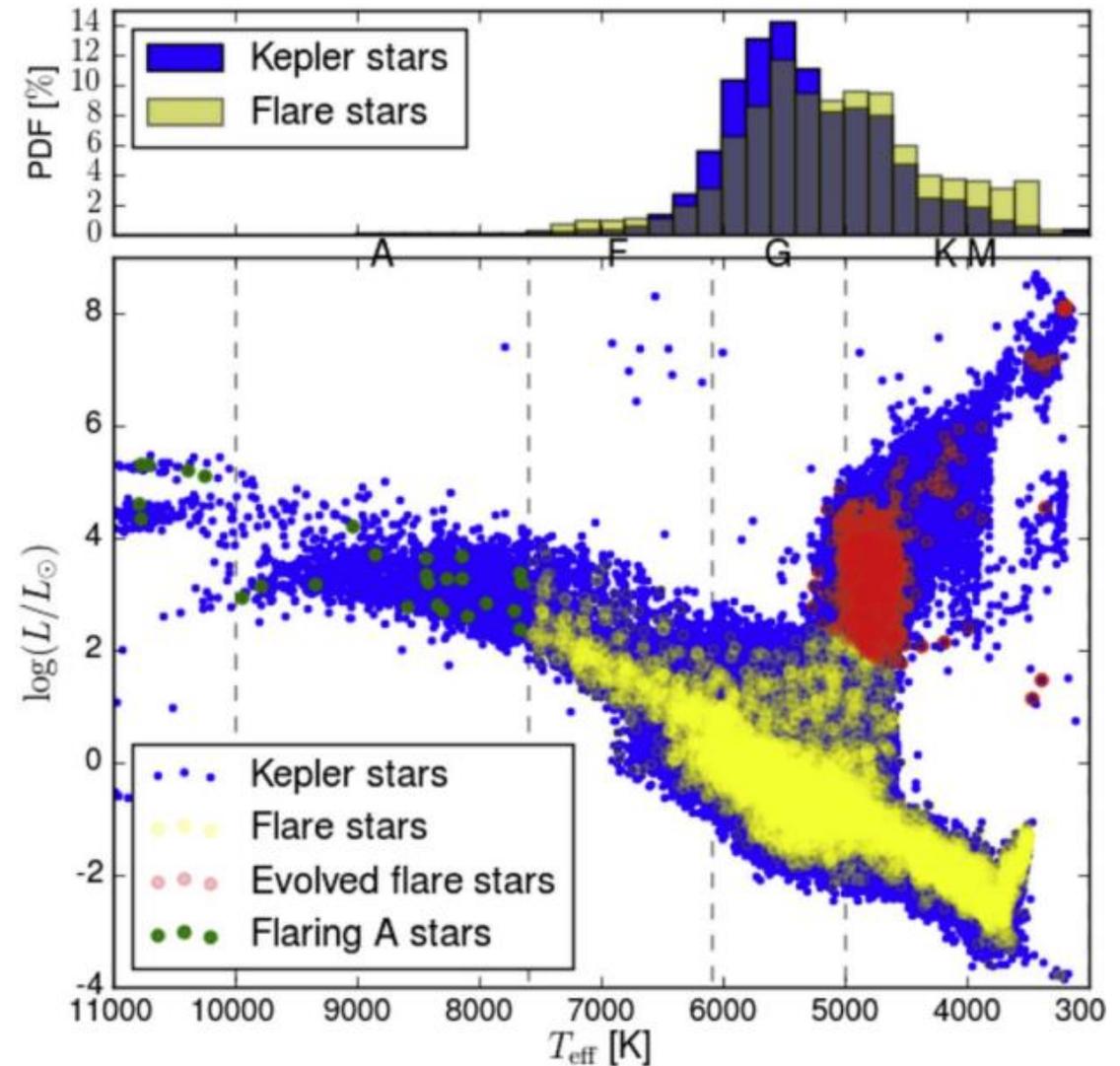


e.g., **H α activity**-Rotation relation

Statistical Properties of Stellar Magnetic Activity?

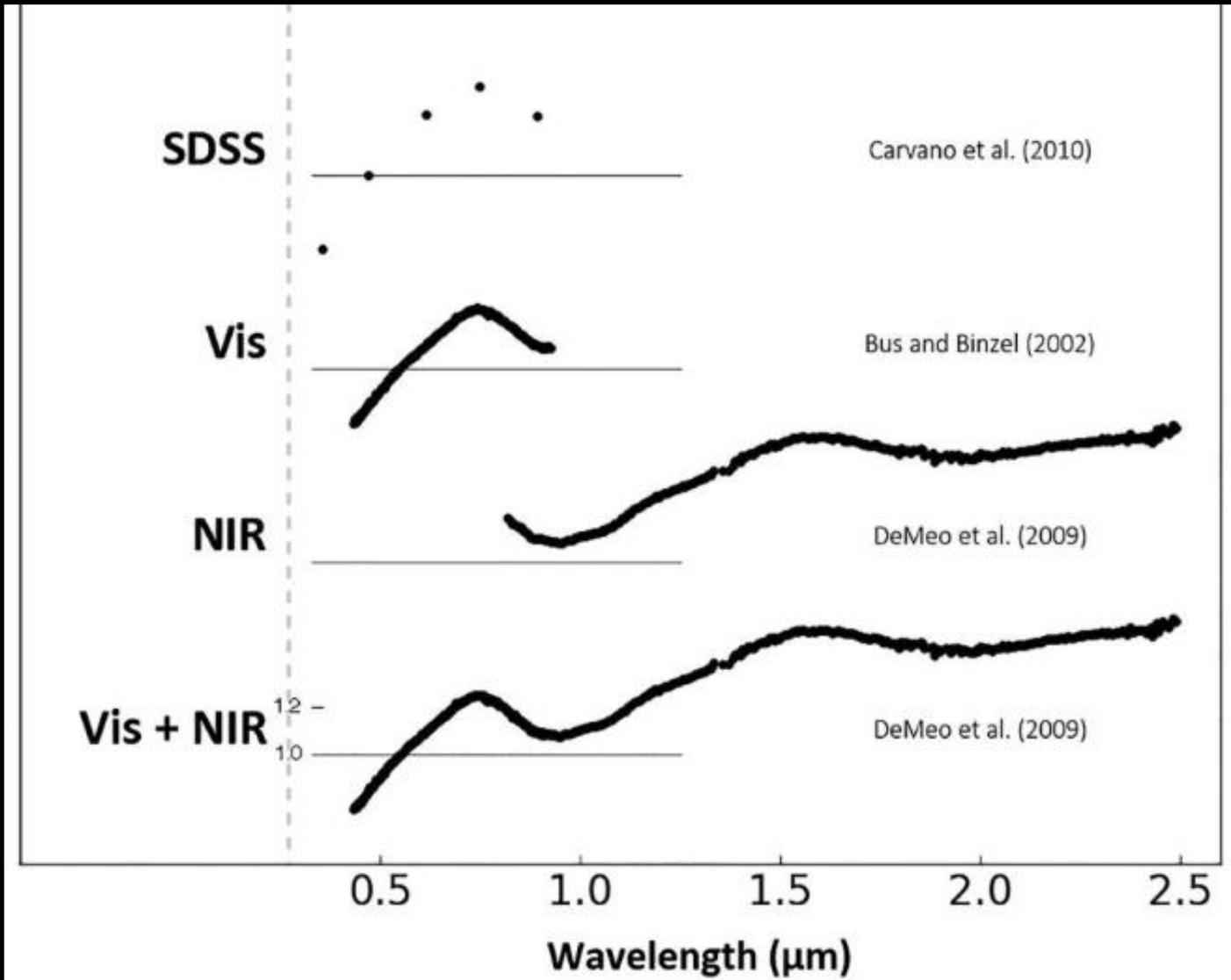


Solar-type stars



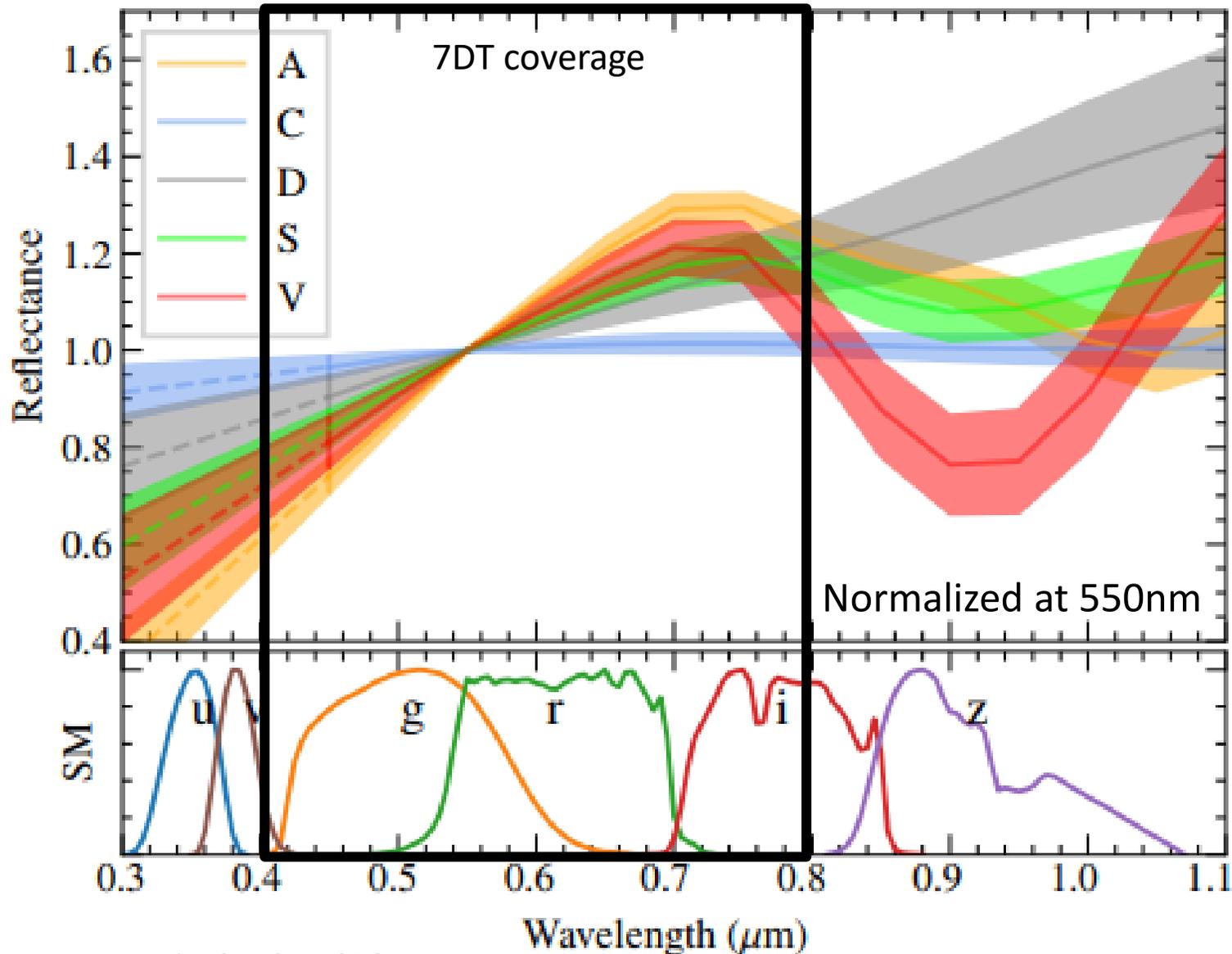
AFGKM + even evolved ones

Latest Galactic Sciences: II. Compositional mapping of the asteroids



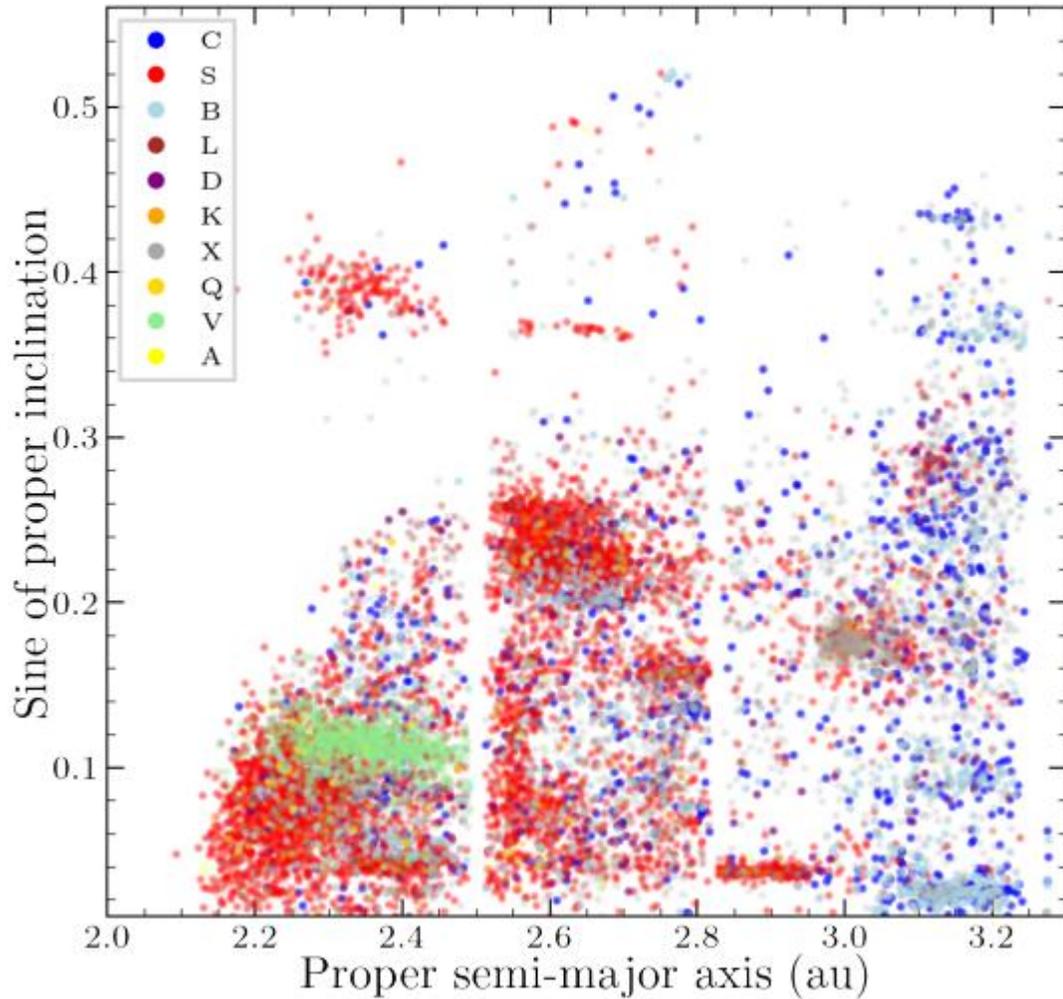
- ~9,000 asteroid spectra: just a tip of the iceberg
- **Multi-filter photometry:** provides strong constraints on asteroid composition and classification, **providing less details but on large samples**
- **SDSS photometry** has been the source of a variety of studies for almost two decades (e.g., dynamic families and surface aging)

Advantage of SkyMapper multi-filter observations

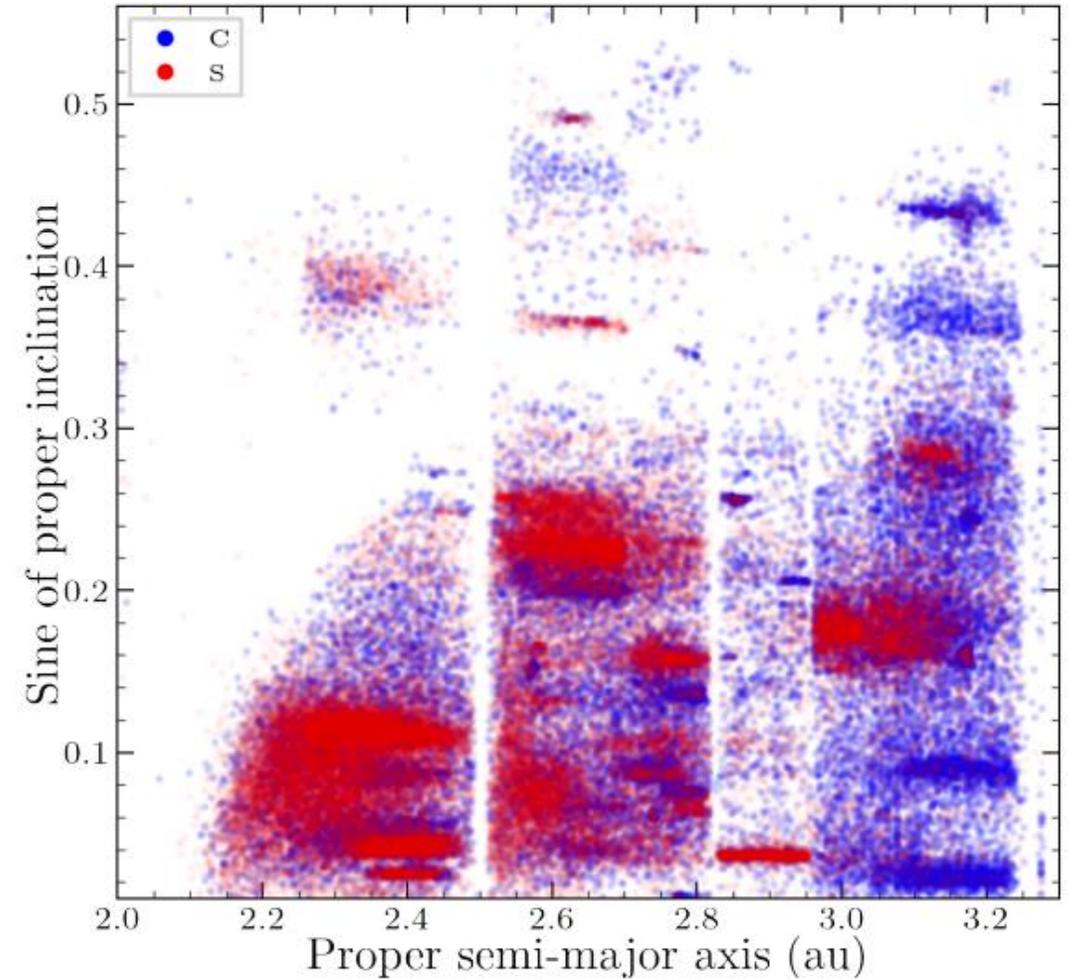
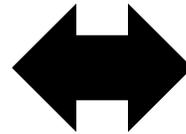


- Spectral degeneration of asteroid spectral classes over visible only (see 7DT coverage)
- **Near-simultaneous acquisition of different filters**
→ **Instantaneous determination of colours.**
- z filter probes the 1 μm band
→ the most characteristic spectral feature in all major taxonomies

Potential of multi-filter taxonomic classification: Dynamical populations?



[g-r-i-z] colours

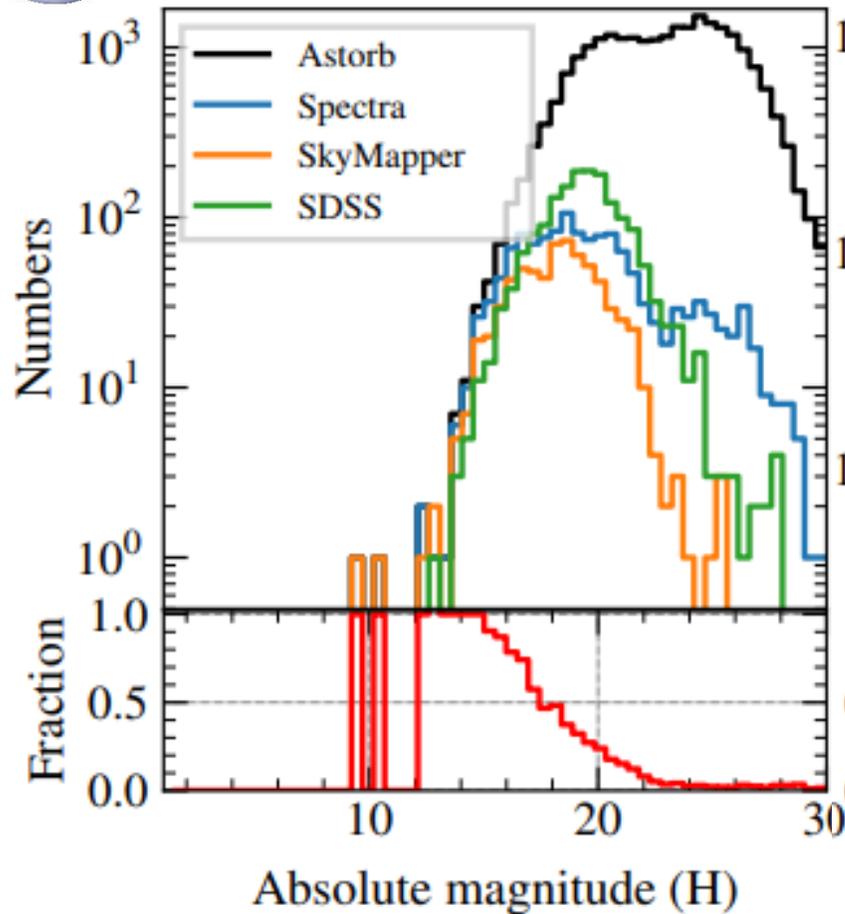


[g-r] colour only (7DT coverage)

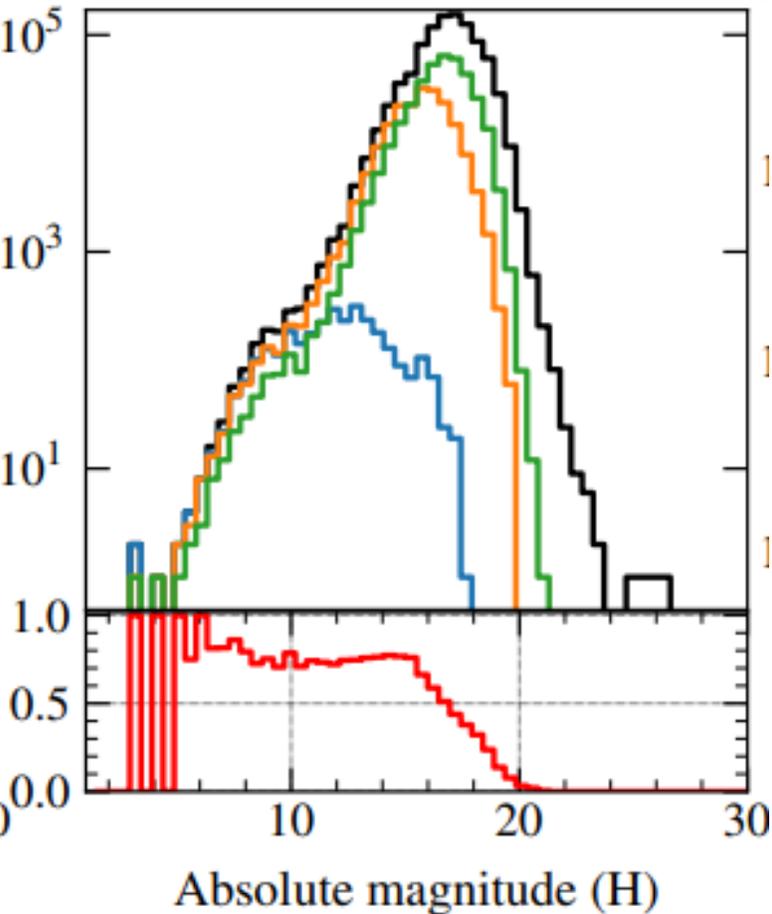
Move toward completeness of taxonomic classification?



Near Earth Asteroid



Main Belt Asteroid



- DR3 205,515 known solar system objects → only a half of them observed by SDSS
- SkyMapper + SDSS → completeness almost 80% of all MBAs with an absolute magnitude between 10 and 16

Last slide: Recommendations for 7DT Galactic Time-Domain Science

The 7DT data will form a time-dependent, multi-colour movie of Southern sky.

- Due to the step-by-step construction of 7DT, **the filter sequence (or filter prioritization: blue to red) and the distribution of revisit interval** are essential to characterizing variability on various timescales from a single survey.

- All 7DT Galactic **Time-Domain sciences are enhanced with Near-simultaneous (< 4-5 minutes), Multi-filter measurements** as seen in the SkyMapper:

 - For stellar activity studies: most commonly **hydrogen Balmer emission lines and continuum** around them.

 - For small body studies: still benefit from a visible only coverage.